## Young alpha rich stars in the solar neighbourhood (MNRAS in press, arXiv:1412.3453)

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## alpha rich stars are expected to be old

Nucleosynthesis in Type II vs Type Ia supernovae



## alpha rich stars are expected to be old

- Nucleosynthesis in Type II vs Type Ia supernovae
- alpha rich stars formed before Type la supernovae could enrich the ISM in Fe





Age [Gyr]



# 1600 stars in the APOKASC sample (APOGEE + Kepler)



## Seismic scaling relations

$$\begin{bmatrix} \nu_{\rm max} \propto g T_{\rm eff}^{-1/2} \propto M R^{-2} T_{\rm eff}^{-1/2} \\ \Delta \nu \propto \rho^{1/2} \propto M^{1/2} R^{-3/2} \end{bmatrix}$$

$$M = \left(\frac{\nu_{\max}}{\nu_{\max,\odot}}\right)^3 \left(\frac{\Delta\nu}{\Delta\nu_{\odot}}\right)^4 \left(\frac{T_{\text{eff}}}{T_{\text{eff},\odot}}\right)^{1.5}$$

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## From mass to age



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- Visual inspection of spectra and light curves: OK
- No anomalous surface rotation, low radial velocity scatter
- Spatial and velocity distribution typical of other alpha rich stars



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Minchev, Chiappini & Martig 2013

## For RC stars, young ages consistent with isochrones



## For RC stars, young ages consistent with isochrones



# Other examples of alpha-rich young stars – (1) CoRoGEE







# Other examples of alpha-rich young stars – (4) galactic center



## Conclusion

- Several surveys find alpha-enriched young stars
- Not predicted by standard chemical evolution models
- Inhomogeneous mixing in the ISM + radial migration

