

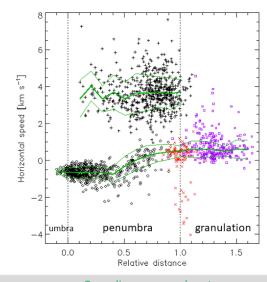


Mysterious motion of penumbral grains

M. Sobotka, J. Jurčák, and M. García Rivas

Astronomical Institute of the Czech Academy of Sciences, Ondřejov, Czech Republic

- Horizontal speeds vs. position Sobotka & Puschmann 2022, A&A, 622, A13
- ♦ penumbral grains (PGs), + dark bodies of filaments, x – penumbral border, □ – G-band bright points
- PGs: the mean speed increases gradually with relative distance d from -0.7 km/s inwards ($d \le 0.4$) to 0.4 km/s outwards (d = 0.8). The direction of PGs motion changes at $d \approx 0.6$
- Why do PGs move inwards in the inner penumbra and outwards in the outer penumbra?



Green lines: mean values ± σ

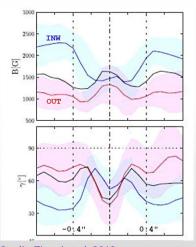
2 Filament model (Tiwari et al. 2013, A&A, 557, A25): transversal cuts of PGs in *B* and γ

The inclination γ of background magnetic field increases gradually with the distance from the umbra, so that:

- Inner penumbra: $\gamma(PG) > \gamma(surroundings)$
- Outer penumbra: γ(PG) < γ(surroundings)

The inclination of surrounding field might affect ascending flows thus the apparent motions of penumbral grains.

Is it really the case? Do inward-moving PGs have their magnetic inclination larger and outward-moving PGs smaller than that in the surroundings?



Credit: Tiwari et al. 2013

AR 13014: Magnetic inclination in PGs and surrounding field

Inclination map with PGs represented by short (0.6") line segments directed along the local magnetic azimuth (green: inwards, 80 PGs; red: outwards, 44 PGs) →

Mean inclination along the PG lines and mean inclinations along two parallel lines (white) on opposite sides of the PG line at distance 0.5" are compared and classified:

- class -1: γ(PG) < γ(both sides) "U"
- class 0: all other cases (unresolved)
- class 1 : γ(PG) > γ(both sides) "∩"

Populations of PGs in the classes \rightarrow

Inwards: most frequent $\gamma(PG) > \gamma(surr.)$

Outwards: most frequent $\gamma(PG) < \gamma(surr.)$

