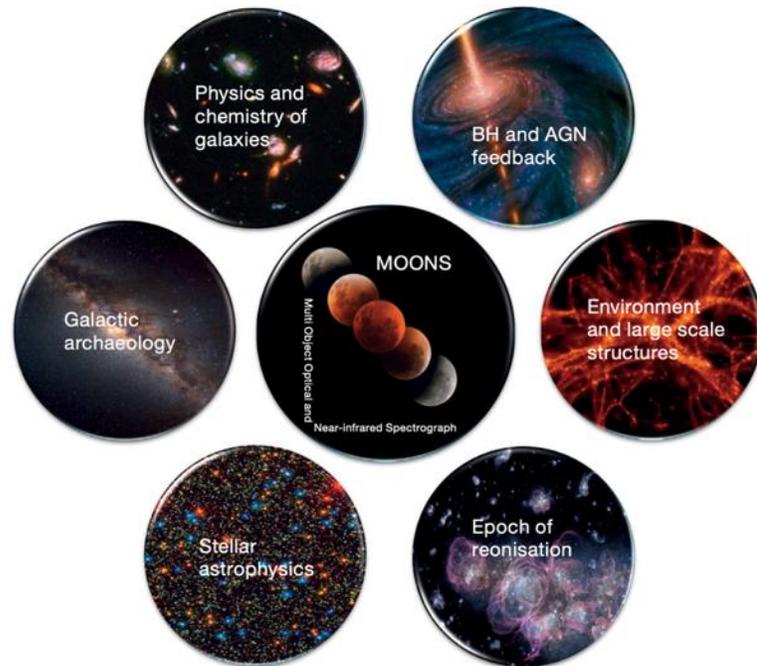




# How 4MOST and MOONS will improve constraints on the dynamical masses of Local Group Galaxies

Steffen Mieske (ESO Chile), on behalf of many colleagues\*

vltmoons.org



4most.eu



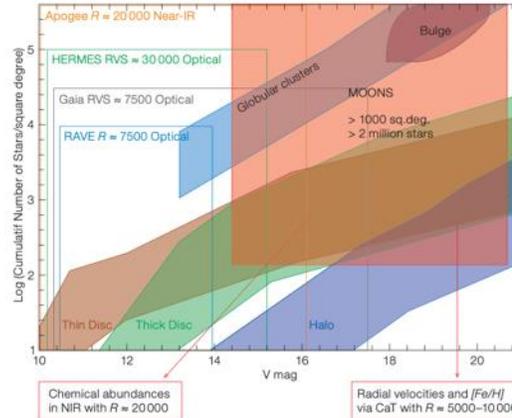
\*Battaglia, Bayo, Cioni, Gonzalez, Mainieri, Mucciarelli, Schmidtobreck, Skuladottir, Szeifert

# MOONS and 4MOST in a nutshell



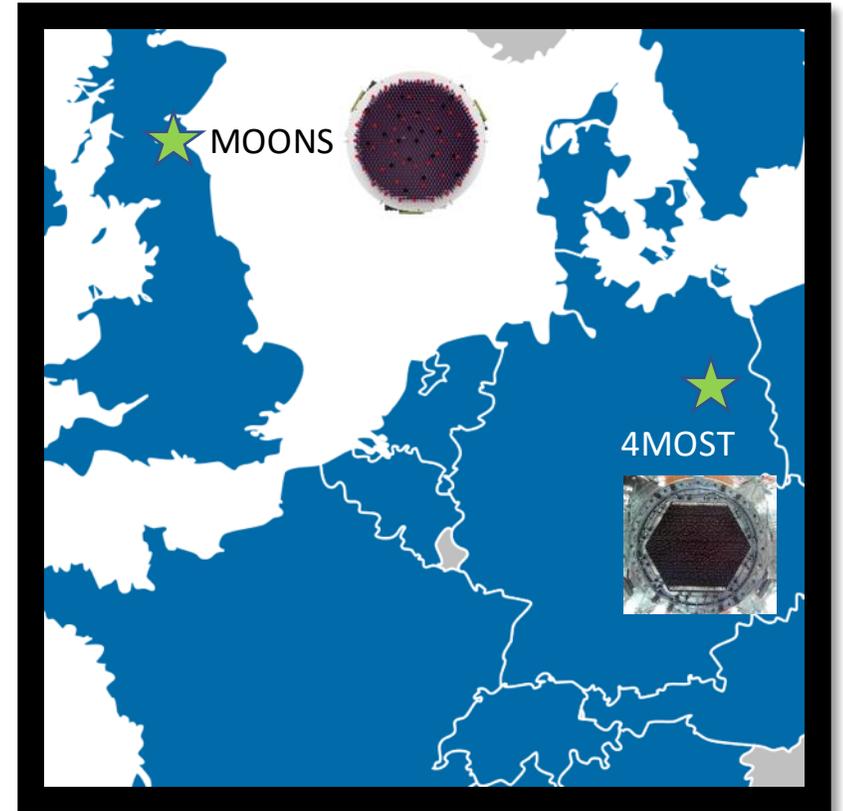
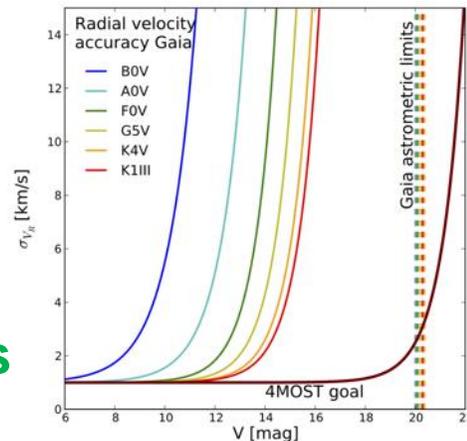
## MOONS:

- MOS@UT1, 1000 fibers, 25'  $\emptyset$
- 6500-18000 Å
- R~4000-20000
- $V < \sim 21$  mag for accurate  $v_{\text{rad}}$  (1hr)
- Start of operations: end of 2024
- 300 GTO nights & **open time**
- PI Cirasuolo (ESO)



## 4MOST:

- MOS@VISTA, 2400 fibers, 2.5°  $\emptyset$
- 3600-9500 Å
- R~4000-18500
- $V < \sim 20$  mag for accurate  $v_{\text{rad}}$  (1hr)
- Start of operations: end of 2024
- **Public surveys during first 5 years**
- PI de Jong (AIP)

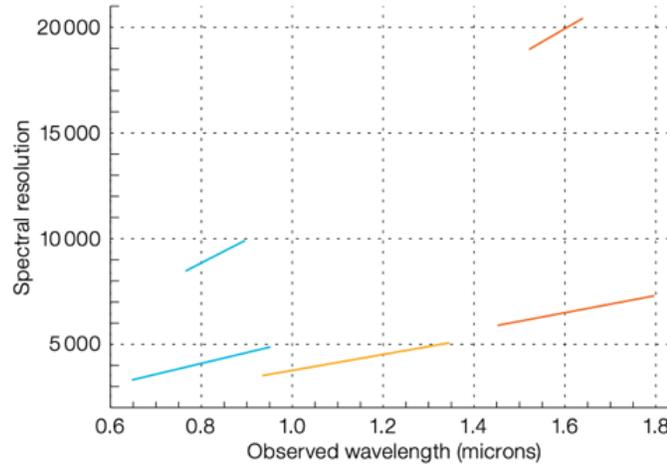


# MOONS and 4MOST in a nutshell

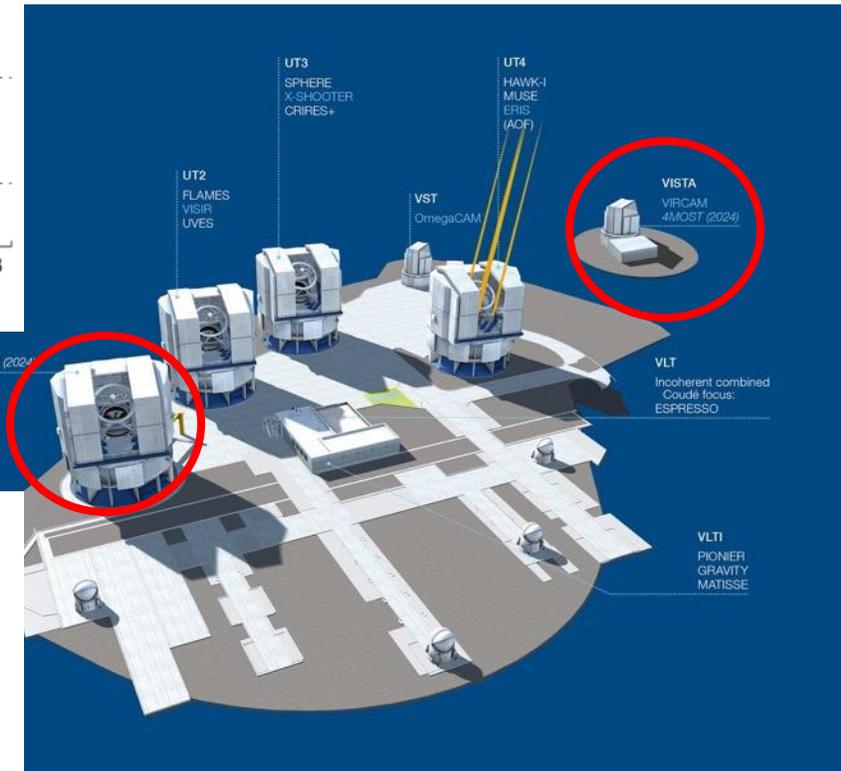


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- MOS@UT1, 1000 fibers, 25'  $\emptyset$
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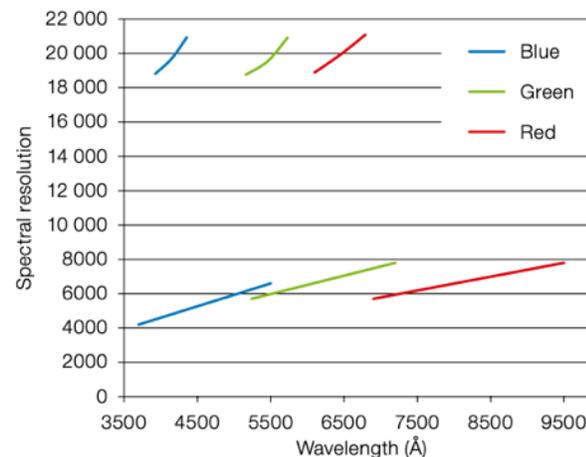


## ESO Paranal



## 4MOST:

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**MOONS FOV 25'  $\emptyset$**   
(full VLT FOV)



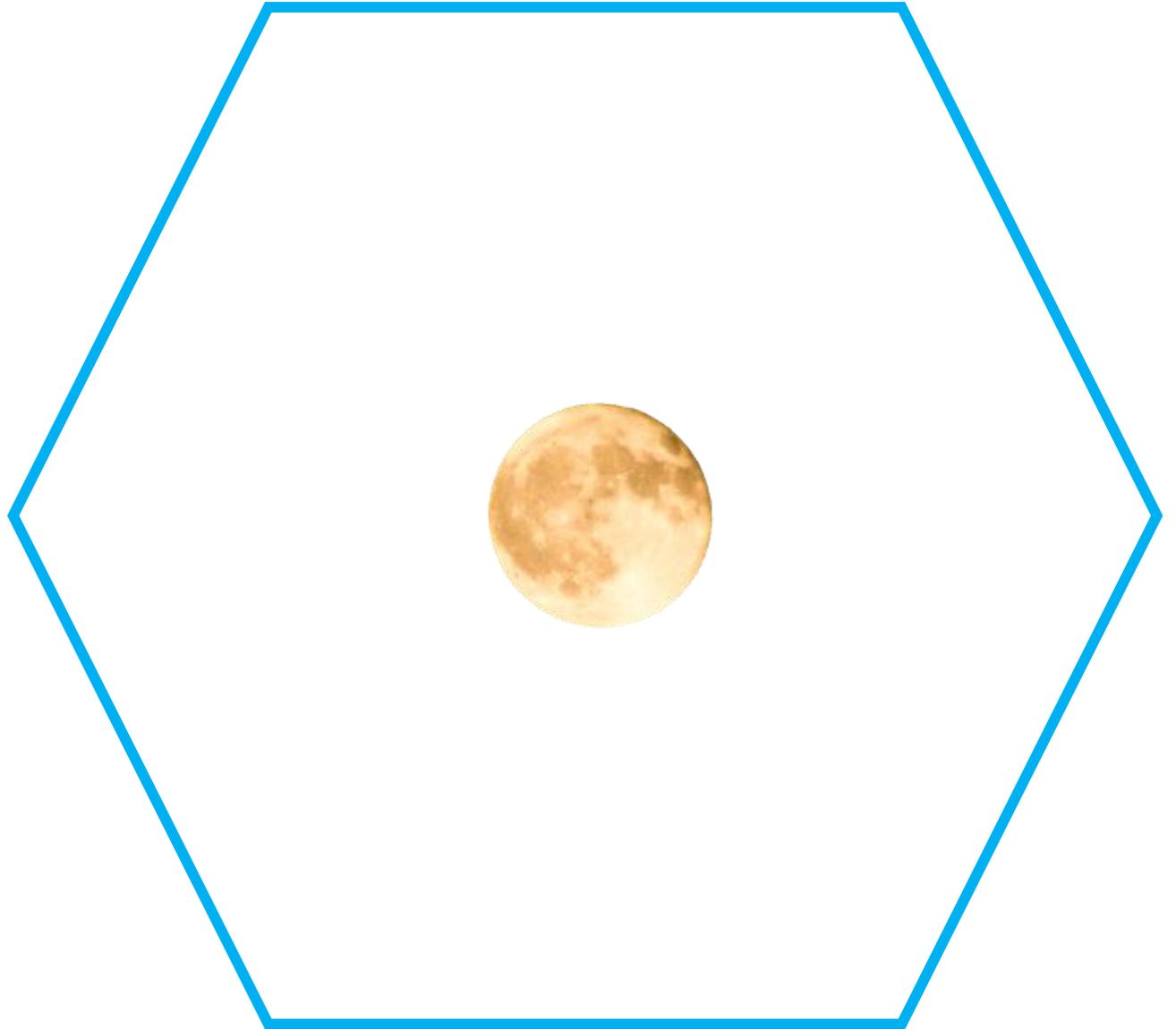
Fornax dSph\*



Sculptor dSph\*

\* ESO DSS2

**4MOST FOV 2.5°  $\emptyset$**   
(VISTA telescope)



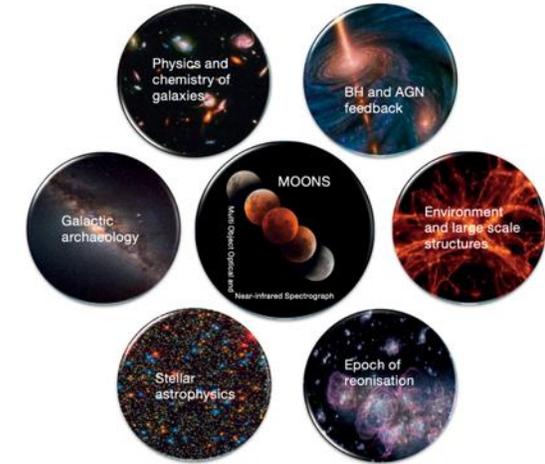
# Science programs with MOONS



## 100 GTO nights Galactic Surveys

- 70 nights bulge
- 30 nights MW satellites (MCs, Sgr, streams)

200 GTO nights Galaxy evolution



**+ Open time**

First Call for Proposals (CfP) expected **2024**





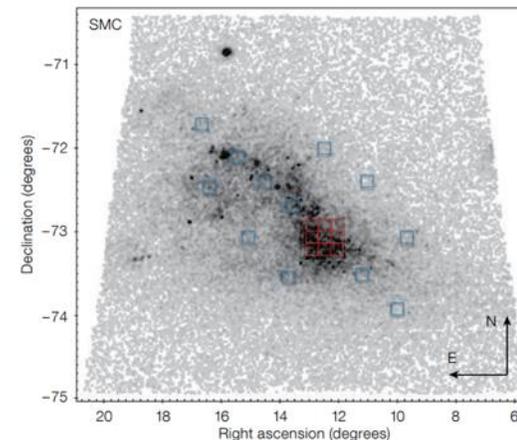
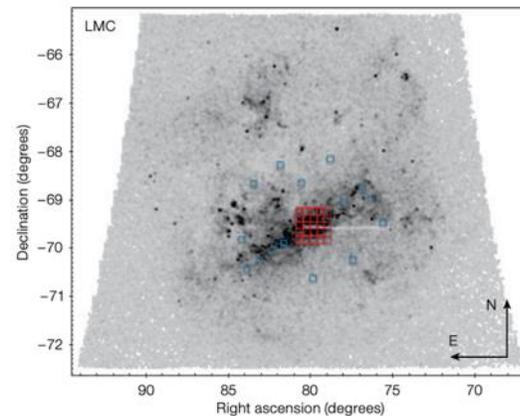
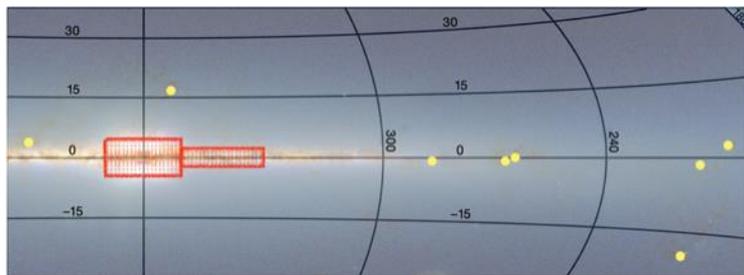
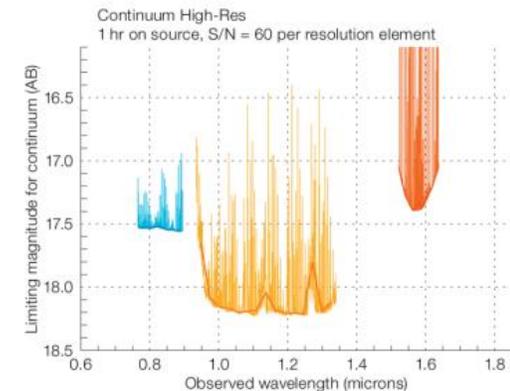
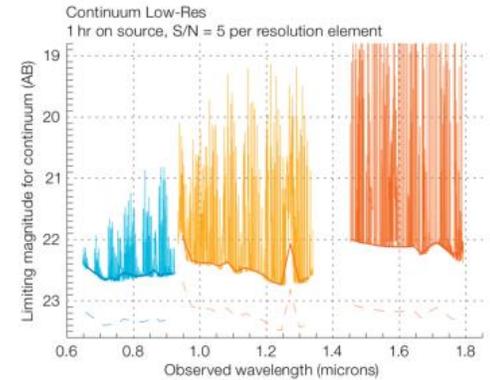
# MOONS Galactic GTO program (100 nights)

## REDdened Milky WAY (REDWAY) survey and Milky Way Satellites Survey

PI: González

### Goals:

- Sample the chemodynamics of the stellar populations in the Galaxy's inner 3 kpc
- Sample the stellar populations in the central regions of the Magellanic clouds and in the Sagittarius galaxy and streams



# Science programs with 4MOST



## Galactic surveys:

Split in community (30% of time) and consortium (70% of time) public surveys

### Stellar Clusters in 4MOST

Survey PI: Sara Lucatello (Osservatorio Astronomico di Padova)

### White Dwarf Binary survey (WDB)

Survey PIs: Odette Toloza (Universidad Técnica Federico Santa María) and Alberto Rebassa (Universitat Politècnica de Catalunya)

### 4MOST Survey of Young Stars (4SYS)

Survey PI: Germano Sacco (Osservatorio Astrofisico di Arcetri)

### 4MOST Gaia RR Lyrae Survey (4GRoundS)

Survey PI: Rodrigo Ibata (Université de Strasbourg)

### Spectroscopic Discovery of Binaries with Dormant Black Holes

Survey PIs: Michał Pawlak (Uniwersytetu Jagiellońskiego w Krakowie) and Tsevi Mazeh (Tel Aviv University)

### 4MOST survey of dwarf galaxies and their stellar streams: Small but fundamental (4DWARFS)

Survey PI: Ása Skúladóttir (Università degli Studi di Firenze)

### Milky Way Halo Low Resolution Survey

Survey PIs: Else Starkenburg (Rijksuniversiteit Groningen), Clare Worley (University of Cambridge)

### Milky Way Halo High Resolution Survey

Survey PI: Norbert Christlieb (Universität Heidelberg)

### Milky Way Bulge and Disk Low Resolution Survey (4MIDABLE-LR)

Survey PIs: Cristina Chiappini (Leibniz-Institut für Astrophysik Potsdam), Ivan Minchev (Leibniz-Institut für Astrophysik Potsdam)

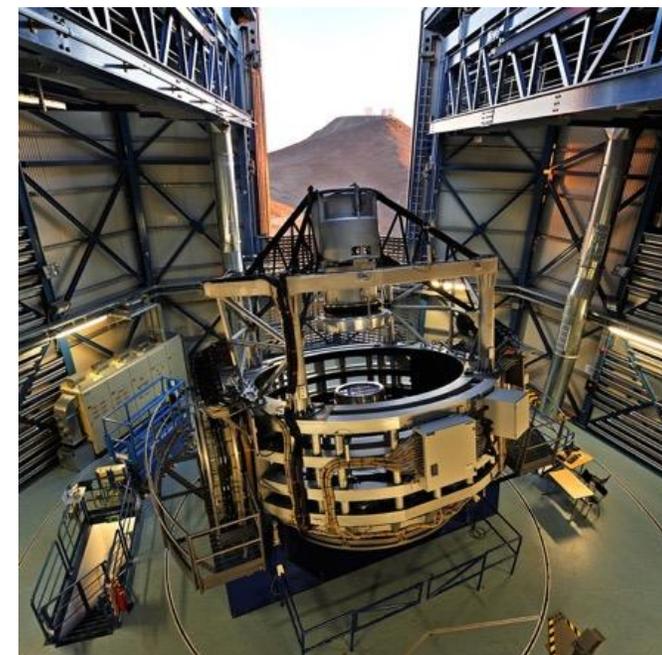
### Milky Way Bulge and Disk High Resolution Survey (4MIDABLE-HR)

Survey PIs: Thomas Bensby (Lunds universitet), Maria Bergemann (Max-Planck-Institut für Astronomie)

### The Thousands and One Magellanic Clouds fields Low and High Resolution Survey (1001MC)

Survey PI: Maria-Rosa Cioni (Leibniz-Institut für Astrophysik Potsdam)

Community  
Galactic



Consortium  
Galactic



4most.eu

# 4MOST surveys aimed at constraining dynamical masses of LG galaxies

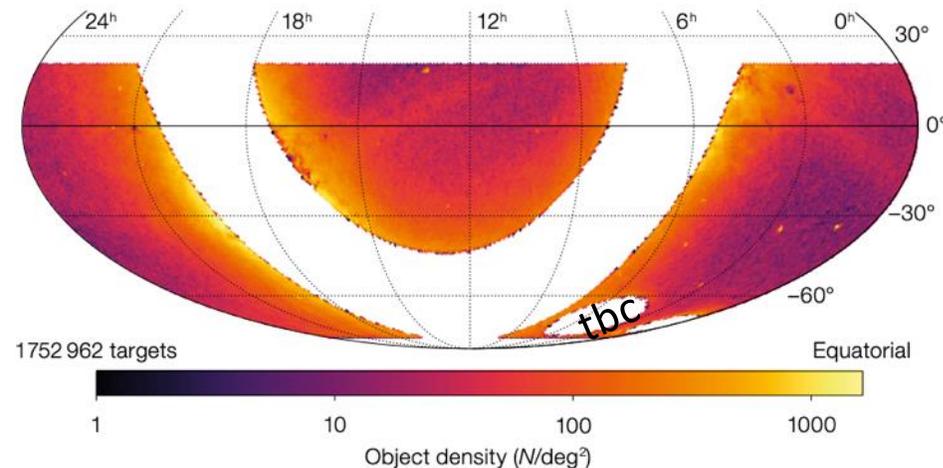


## 4MOST Consortium Survey 1: The Milky-Way Halo Low-Resolution Survey

PIs: Starkenburg & Worley. <https://doi.eso.org/10.18727/0722-6691/5120>

### Goals:

- Determine density profile, shape and characteristic parameters of the DM halo of the Milky Way, including testing alternative theories of gravity such as MOND.
- Measurement of perturbations induced by clumps on the spatial and kinematic properties of cold streams, leading to constraints on the mass spectrum of perturbers and the nature of DM
- Quantifying the amount of kinematic substructure as a function of distance and location on the sky



# 4MOST surveys aimed at constraining dynamical masses of LG galaxies

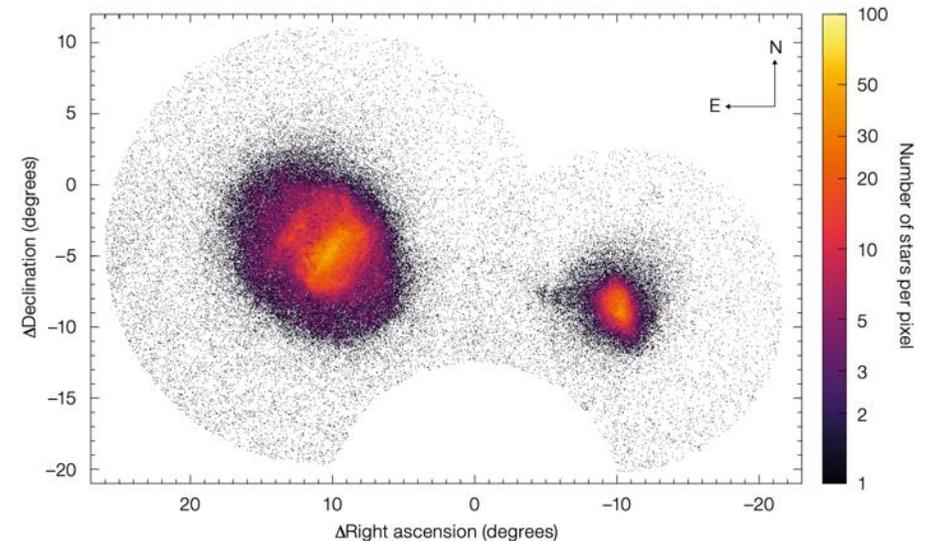
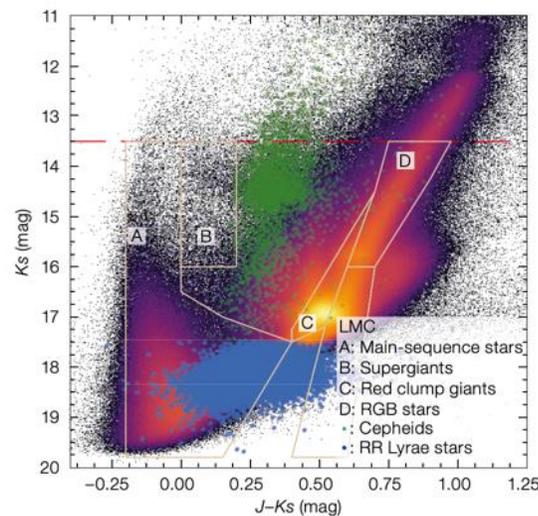
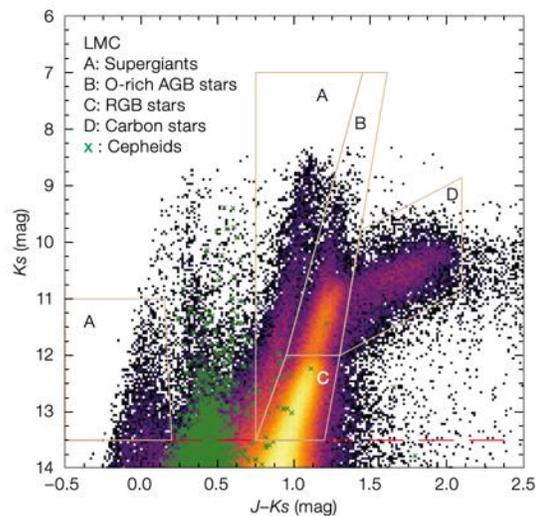


## 4MOST Consortium Survey 5: The Thousand and One Magellanic Cloud Fields Low- and High-resolution Survey (1001MC)

PI: Cioni <http://www.eso.org/sci/publications/messenger/archive/no.175-mar19/messenger-no175-54-57.pdf>

### Goals:

- Identify chemical and kinematical patterns in the Magellanic Clouds, and their interrelations
- Establish how star formation history and dynamical evolution of the MCs are linked to those patterns
- Quantify metallicity dependence of key distance indicators
- Study physical and wind properties of massive stars



# 4MOST surveys aimed at constraining dynamical masses of LG galaxies



## 4MOST Community survey 4: [Gaia RRLyrae Survey](#)

PI Ibata

### Goals:

- Reveal the six-dimensional structure of the outer MW
- Derive non-parametric measure of the MW's dark matter distribution in 3D

4MOST Gaia RRLyrae Survey  
(4GRounds)

Anke Arentsen, Giuseppina Battaglia, Michele Bellazzini, Paolo Bianchini, Gisella Clementini, Benoit Famaey, Alessia Garofalo, Vanessa Hill, Rodrigo Ibata, Khyati Malhan, Nicolas Martin, Alessio Mucciarelli, Giacomo Monari, Lorenzo Posti, Arnaud Siebert, Asa Skuladottir, Antonio Sollima, Guillaume Thomas, Chris Wegg, Zhen Yuan

INSU, CNRS, ESO, ERC logos

The poster features a central image of the Milky Way galaxy. At the bottom, there are logos for INSU, CNRS, ESO, and ERC. The text lists the names of the survey members.

Increase by 2 orders of magnitude the number of RRLyrae with measured  $v_{\text{rad}}$

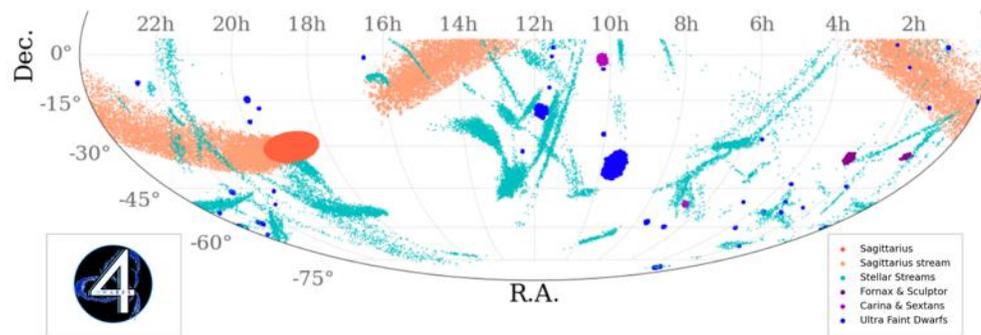
# 4MOST surveys aimed at constraining dynamical masses of LG galaxies



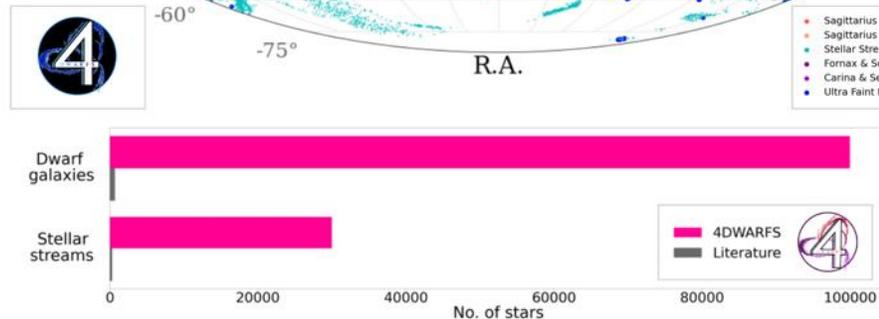
4MOST Community survey 6: [Survey of dwarf galaxies and their stellar streams: small but fundamental \(4DWARFS\)](#). PI Skuladottir

## Goals:

- Constrain dynamical masses
- Identify and study kinematical substructures
- Derive chemo-chrono-dynamical mapping
- Constrain binary distribution from repeat observations



130000 target stars in total



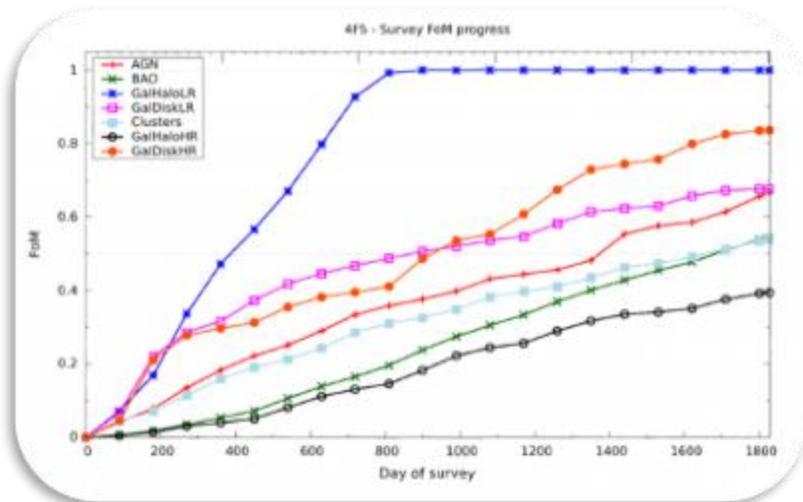
# Community data access



## 4MOST @ VISTA: data products from surveys

The first five years of operations, 4MOST will execute exclusively Public Surveys that have already been selected, see also previous slides.

- **raw data become public immediately** (but no public pipeline available)
- **data products to be released on a ~yearly basis** (spectra, physical parameters, catalogues)
- **MW halo survey could be the first one to be completed**



Preliminary!! –  
<https://www.4most.eu/cms/operations/surveysimulations/>

# Community data access



## MOONS @ UT1: open time proposals

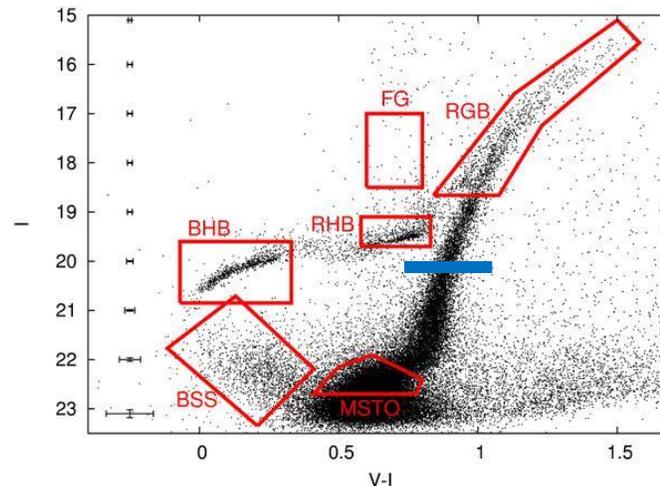
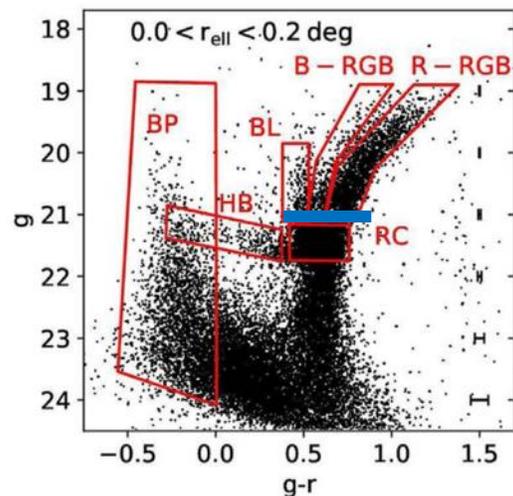
- Offered for observations by end of 2024 via regular CfP
- A public data reduction pipeline will be available



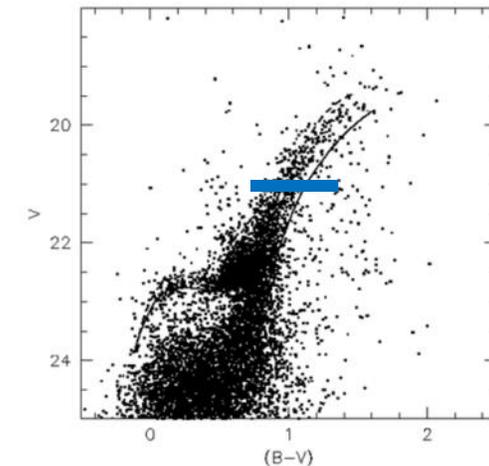
## Kinematical studies of LG dwarfs under-represented in GTO

- Very good stability of MOONS line-spread function
  - MOONS to reach 0.5 km/s in HR, ~2 km/s in LR ( $\sigma$  of LG dwarfs 4-10 km/s)
  - Limiting mag  $V \sim 21$  for  $S/N=20$  in LR (1 mag deeper than GAIA astrometry)

Fornax dSph  
(Wang+ 2019)



Sculptor dSph  
(de Boer+ 2011)

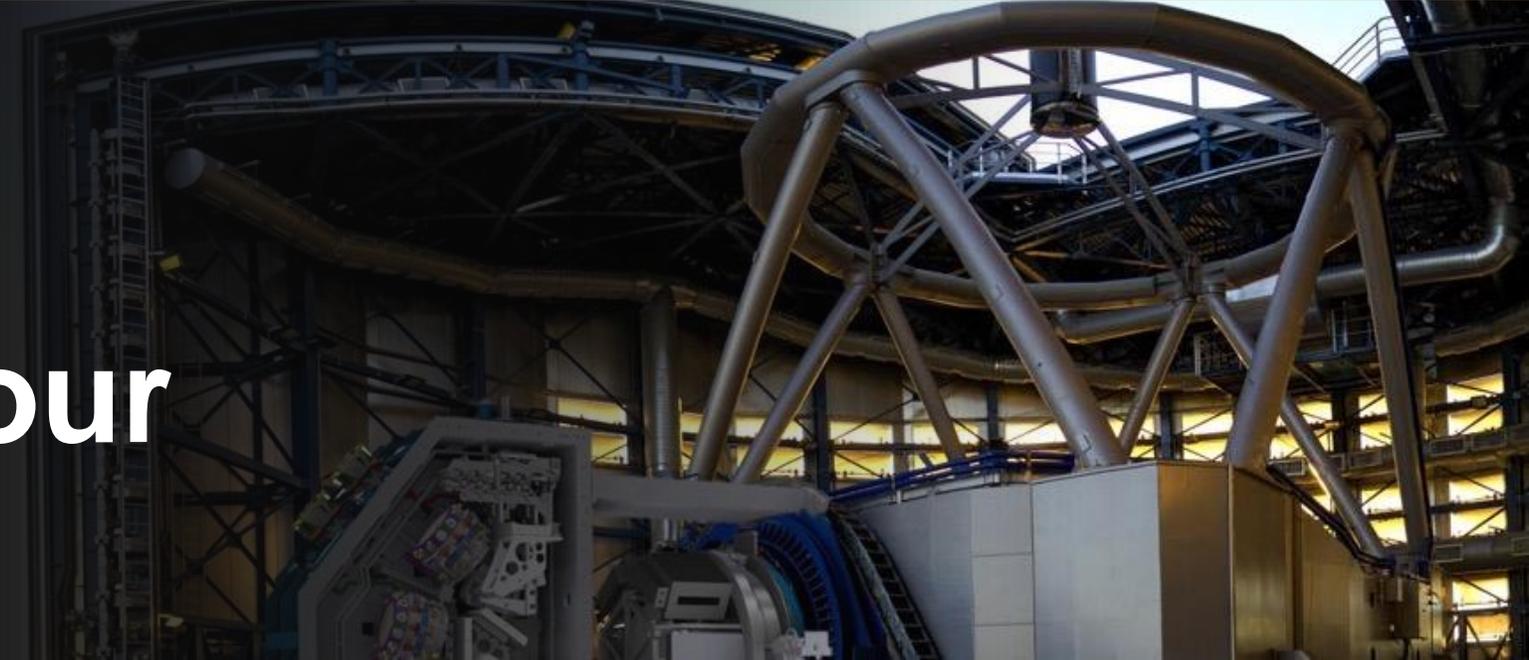
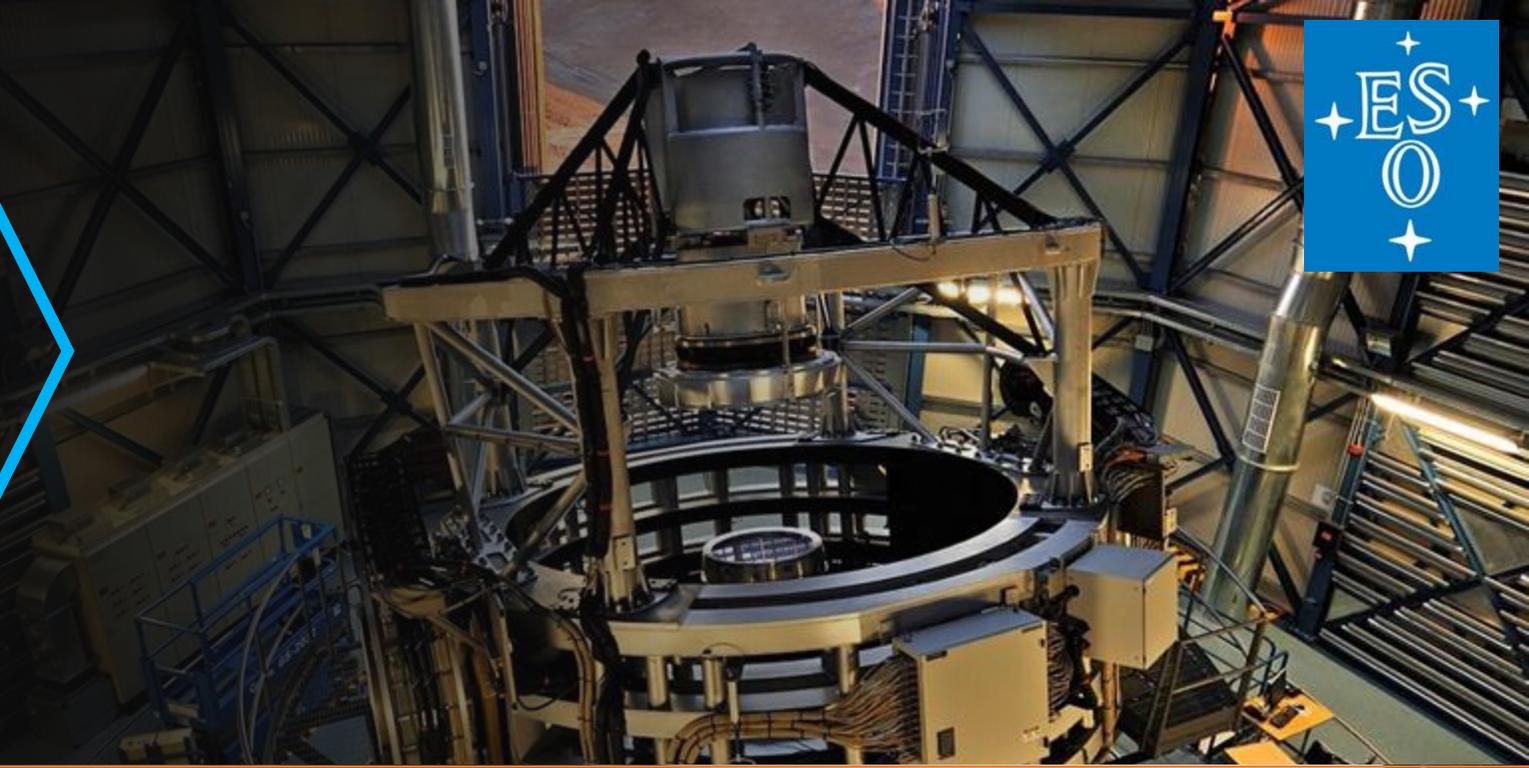


Leo dSph  
(Saviane+ 2000)

# Science questions in the field of this conference that MOONS and 4MOST will help to address



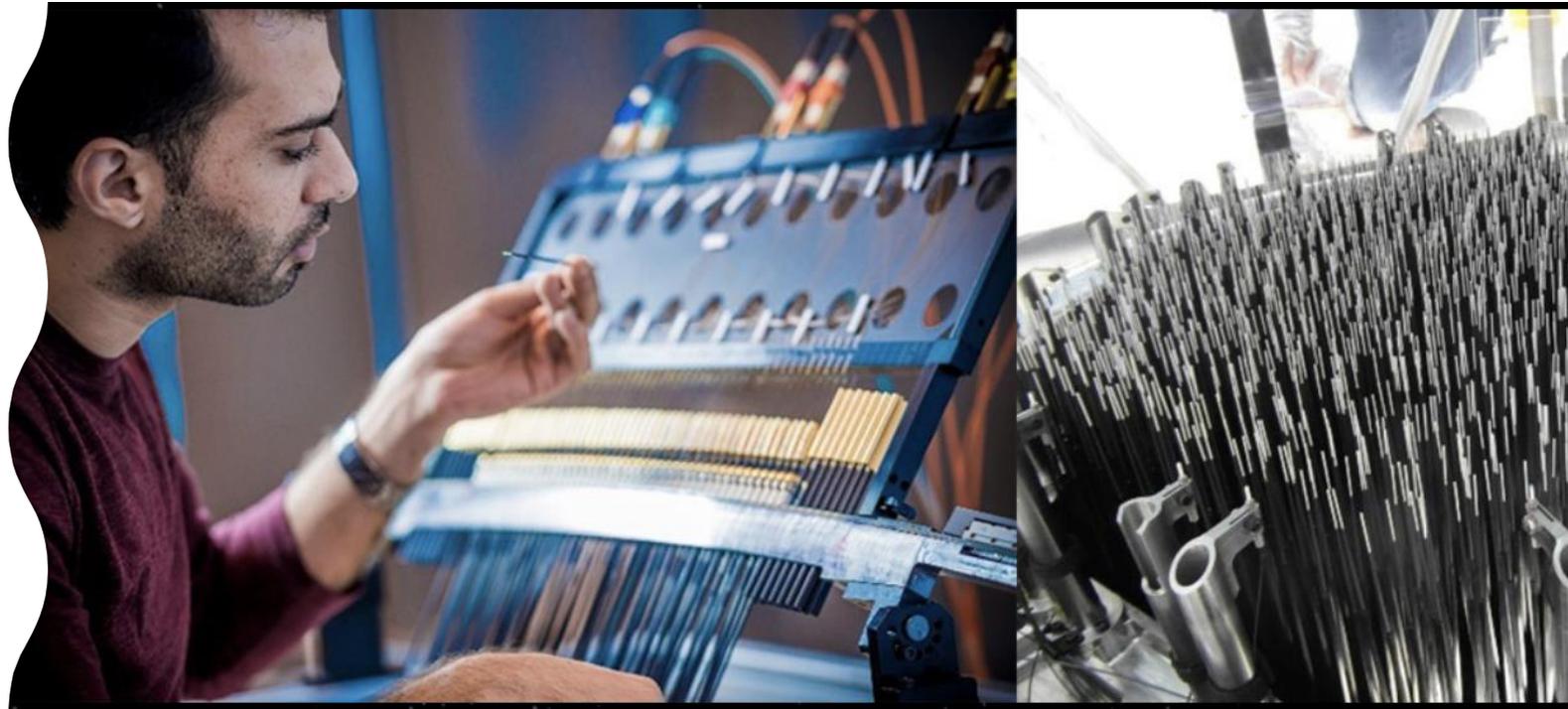
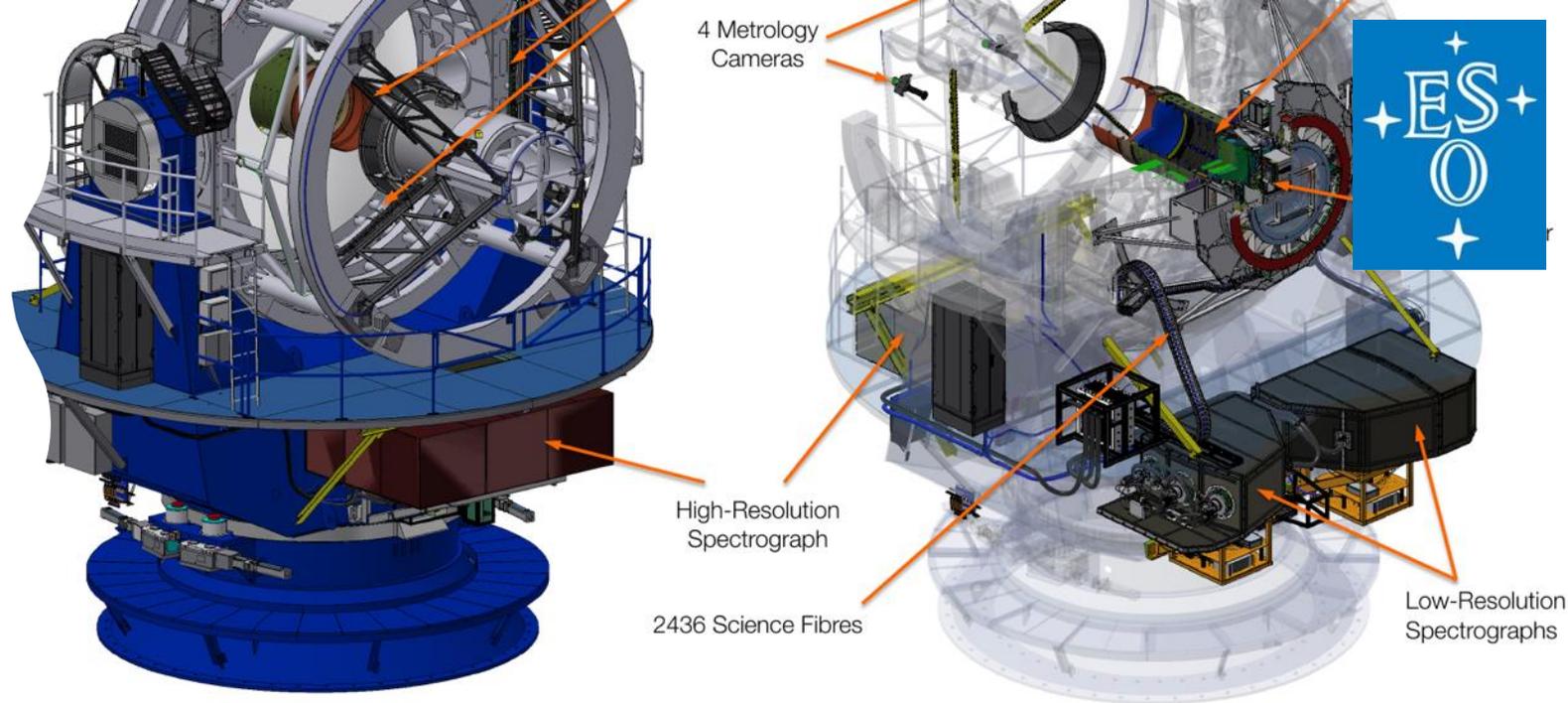
- **Constrain the gravitational potential and dark matter halo of the Milky Way, test alternative gravity theories**
  - 4MOST community survey 4GROUNDS (Ibata), 4MOST consortium survey 1 (Starkenburger)
- **Identify and characterize kinematical and chemical sub-structures of LG dwarf galaxies**
  - 4MOST community survey 4DWARFS (Skuladottir), 4MOST consortium survey 1001MC (Cioni)
  - MOONS GTO: MCs & SgrDwarf (González)
  - **Opportunity for MOONS open time on LG dwarfs!**
- **Characterise properties of stellar streams**



**Thanks for your  
attention**

# 4MOST operations model

- Data for different surveys are taken simultaneously in each exposure
- Instrument consortium takes a service role for the community during the preparation of Observing Blocks and data processing
- Shared responsibility ESO & consortium



# THE EXORCIST

