

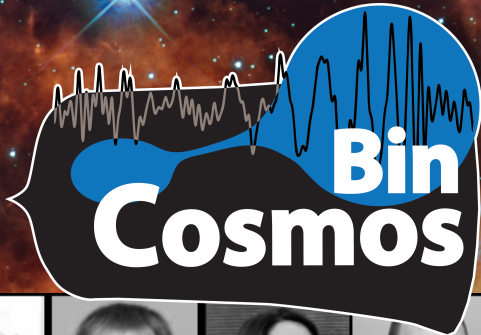













Massive Stars as Galactic Engines

Potsdam Thinkshop on feedback in galaxy formation



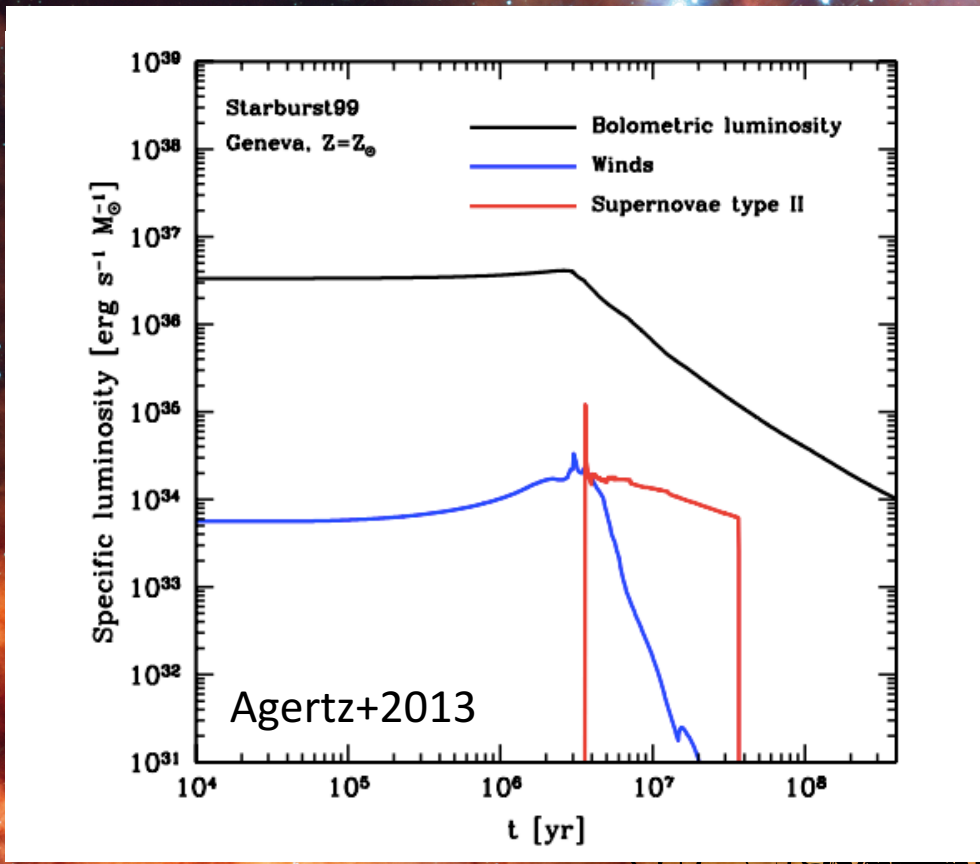
credit: background NASA Peresce, Design: E. Buunk



 Selma de Mink PI	 Stephen Justham Associate '18-	 Silvia Toonen VENI fellow '17-	 Rob Farmer Postdoc '17-	 Athira Menon Postdoc '18 -
 Ylva Göteborg PhD '18	 Manos Zapartas PhD '18	 You?	 Mathieu Renzo PhD '19	 Eva Laplace PhD '21
 Floor Broekgaarden MSc	 Walter van Rossem MSc	 David Hendriks MSc	 Karel Temmink MSc	 Thomas Dodds BSc

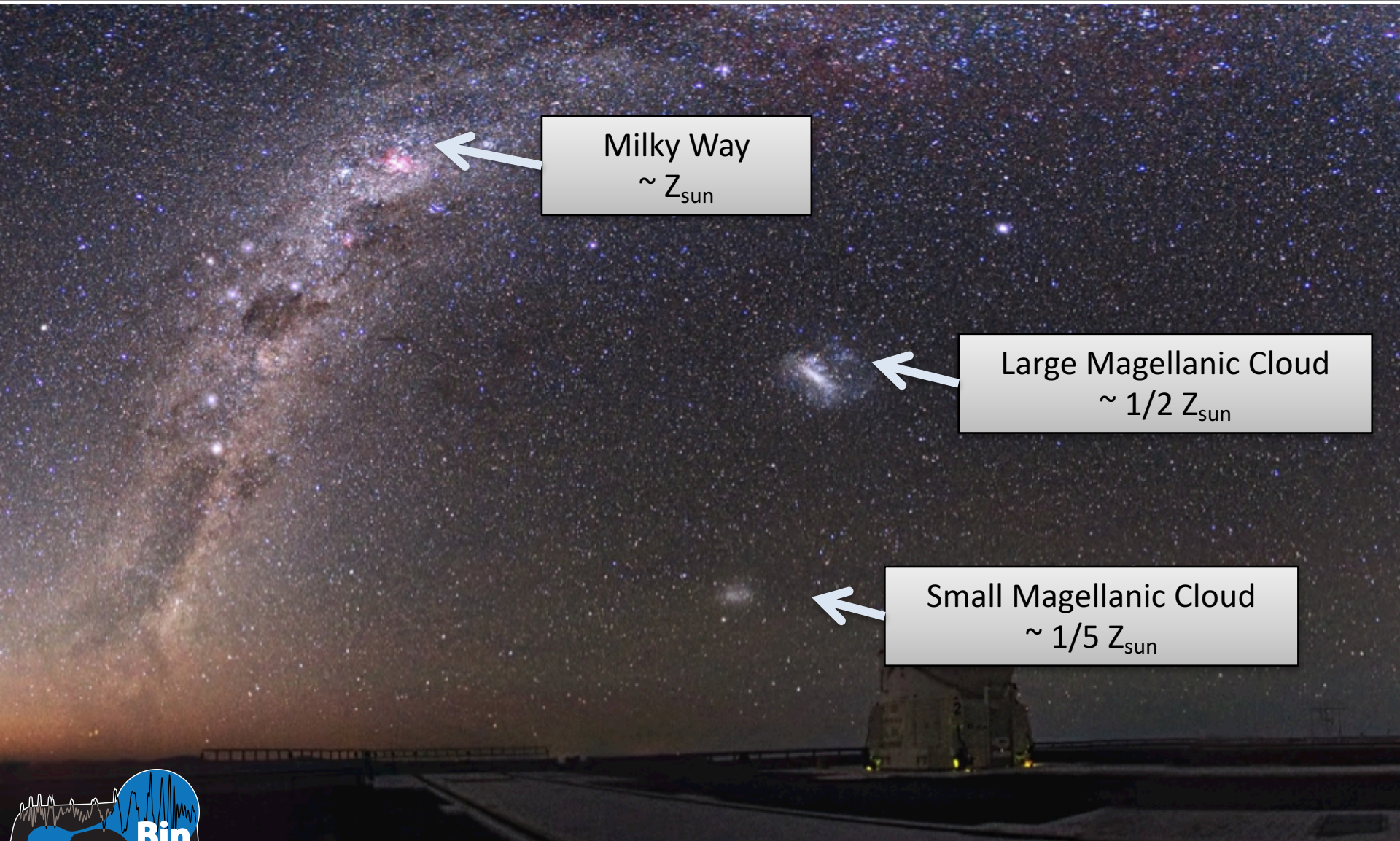


Selma E. de Mink
University of Amsterdam



TOWARD A COMPLETE ACCOUNTING OF ENERGY AND MOMENTUM FROM STELLAR FEEDBACK IN GALAXY FORMATION SIMULATIONS

A massive star Safari

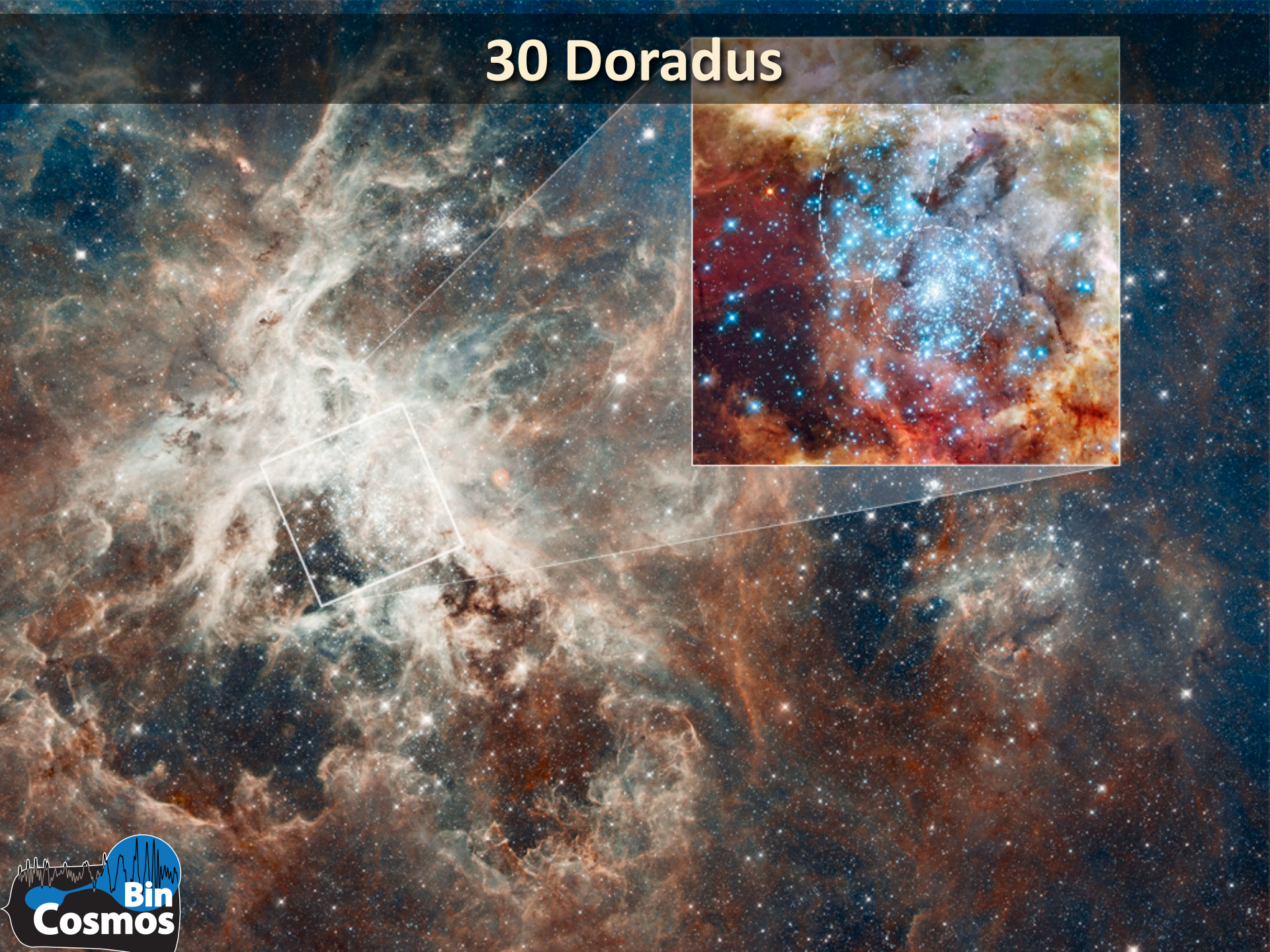


Milky Way
 $\sim Z_{\text{sun}}$

Large Magellanic Cloud
 $\sim 1/2 Z_{\text{sun}}$

Small Magellanic Cloud
 $\sim 1/5 Z_{\text{sun}}$

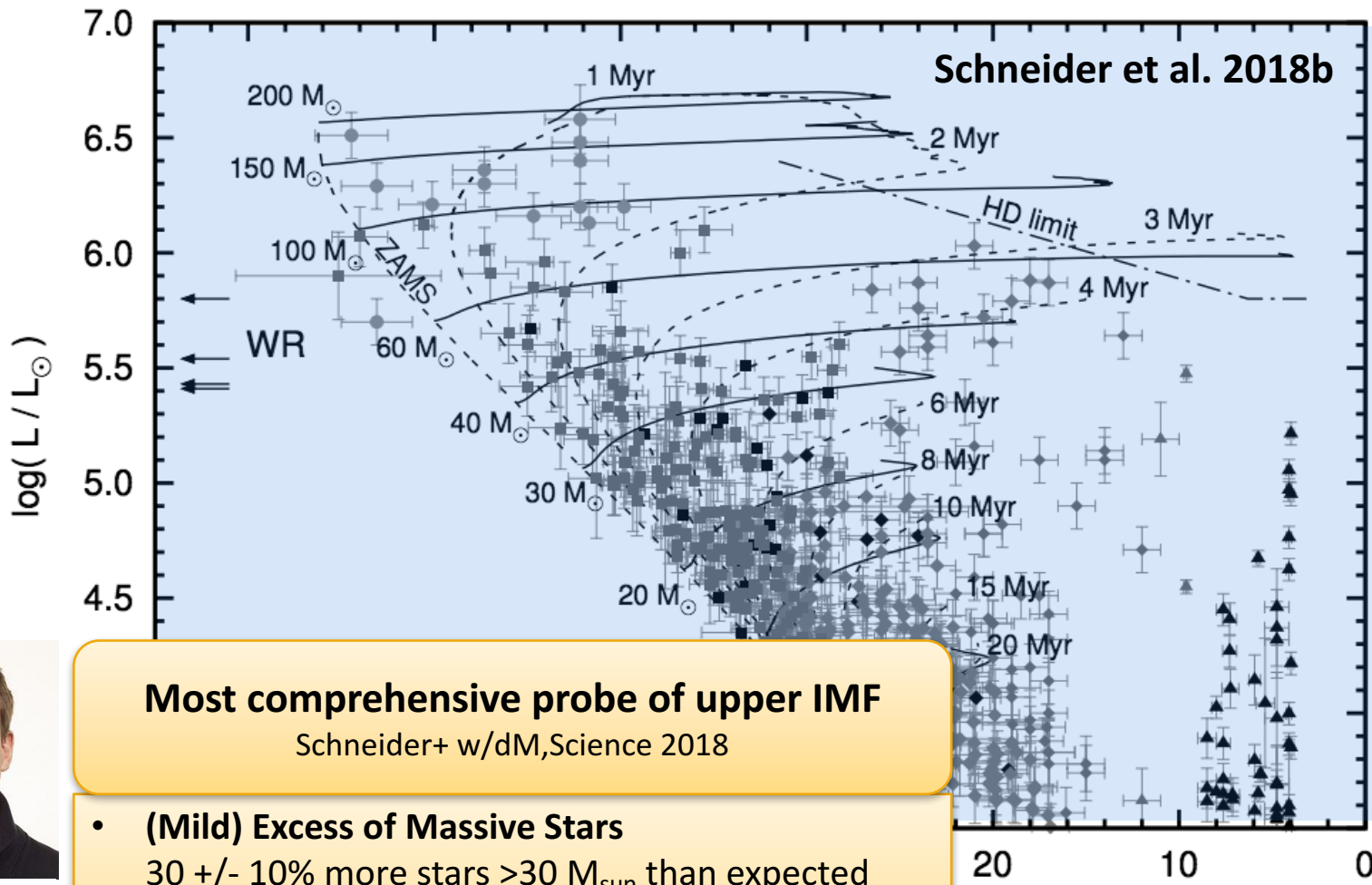
30 Doradus



VLT-FLAMES Tarantula Survey

Evans et al. 2011

multi-epoch optical spectra of over 900 massive stars



Most comprehensive probe of upper IMF

Schneider+ w/dM, Science 2018

- **(Mild) Excess of Massive Stars**
30 +/- 10% more stars >30 M_{sun} than expected
- **Flatter IMF but consistent with Salpeter:**
 $\alpha = 1.9^{+0.37}_{-0.26}$ ($\alpha_{\text{Salpeter}} = 2.35$)

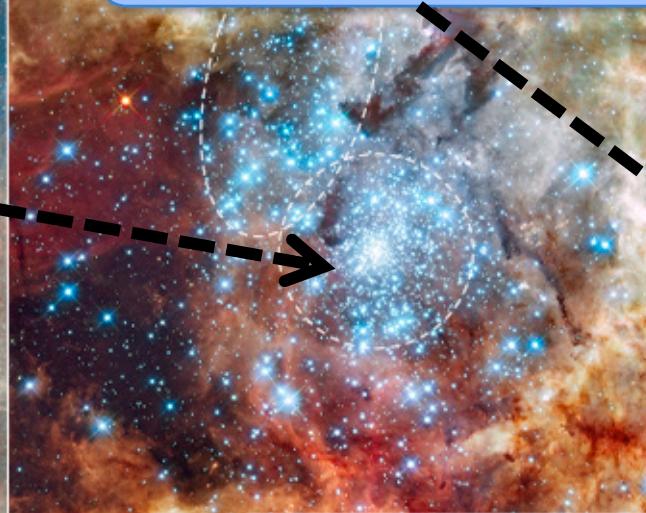


Fabian
Schneider

"Monster Stars" > 100 - 250 M_{sun}
e.g Crowther+ w/dM 2015

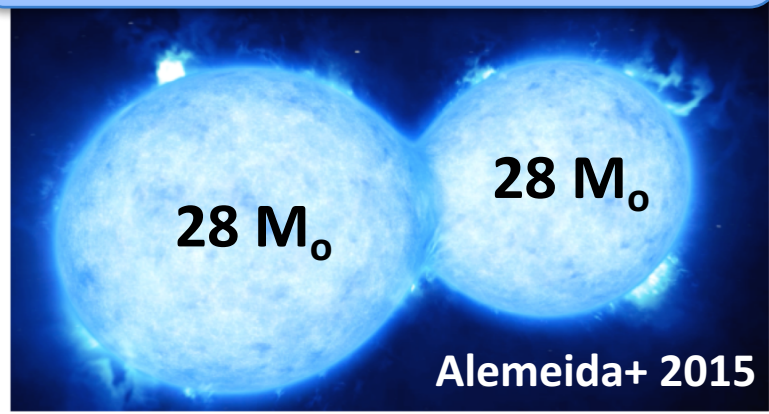
Runaway Stars

Lennon+2018, Renzo+2018b, Sana+prep



Binaries binaries binaries!

Sana+2013, Almeida+2016



Spinning stars

Dufton+w/dM+11,13, Ramirez-agudelo++w/dM13,15

How complete is this picture?

Stars > 100 Msun?

X-ray binaries

Stripped stars in binaries?

Trends with metallicity?

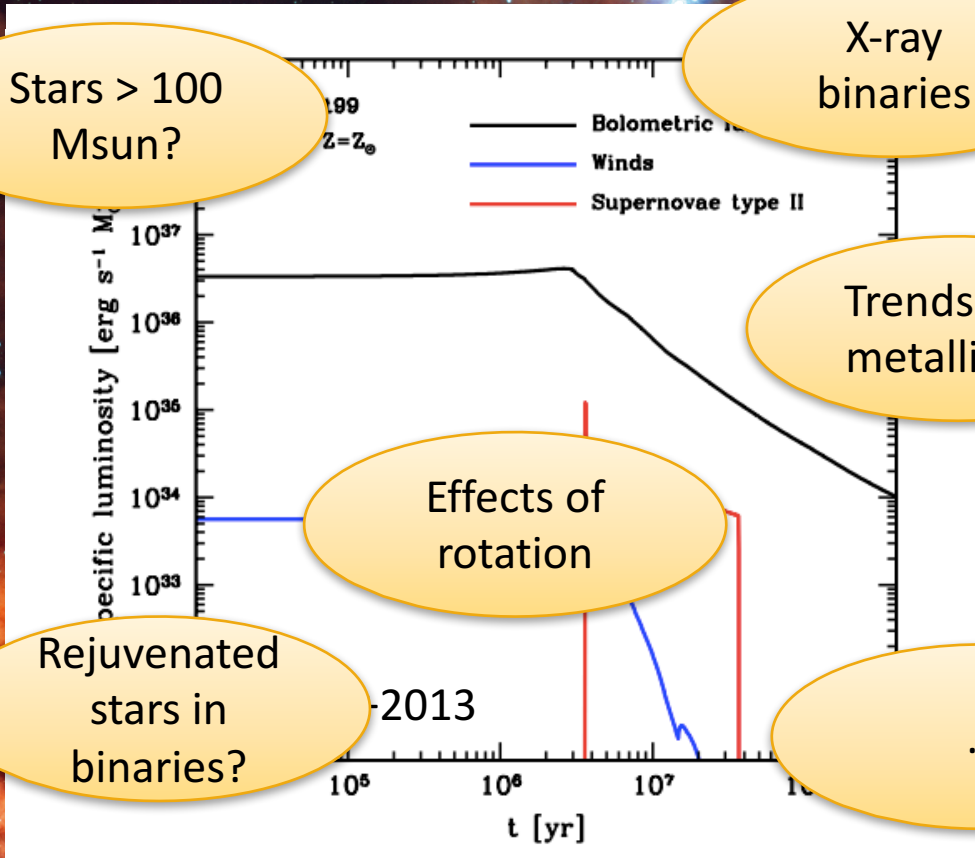
Walkaway / runaways

Effects of rotation

Rejuvenated stars in binaries?

Triples

...

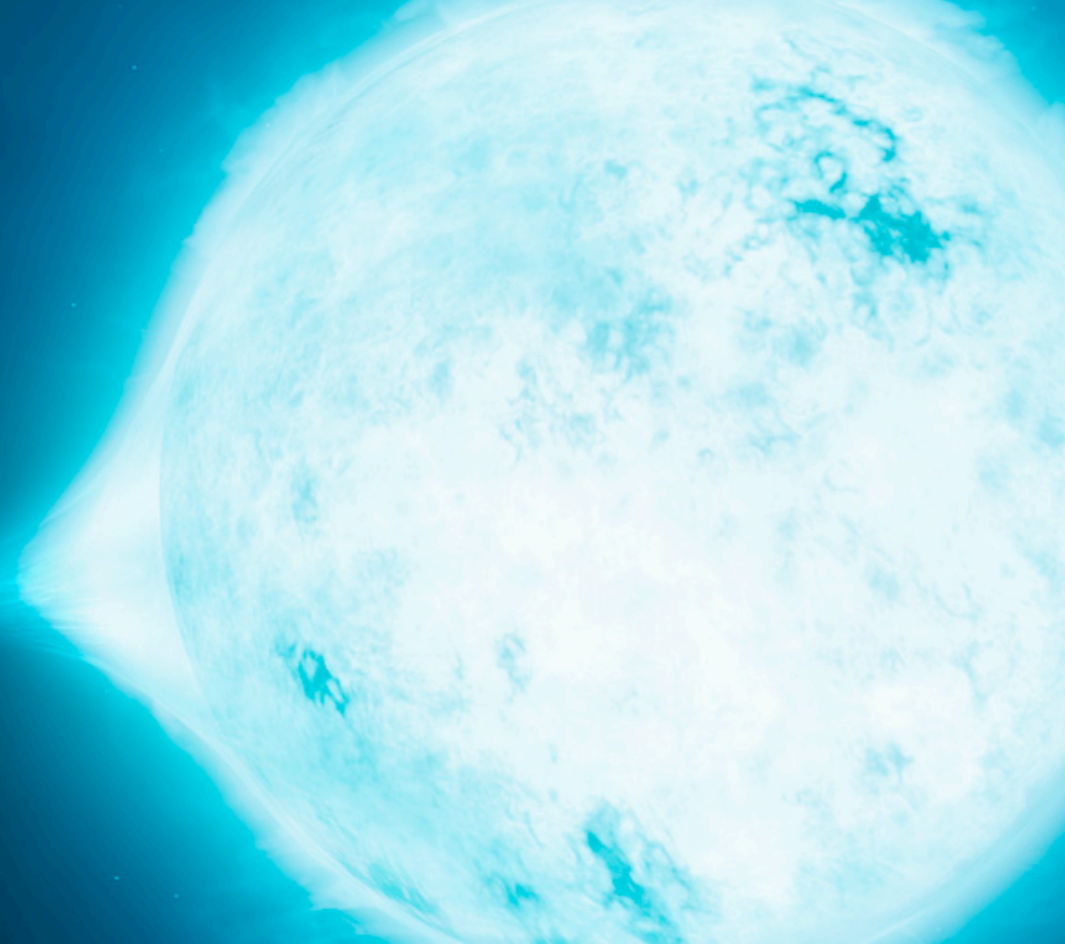
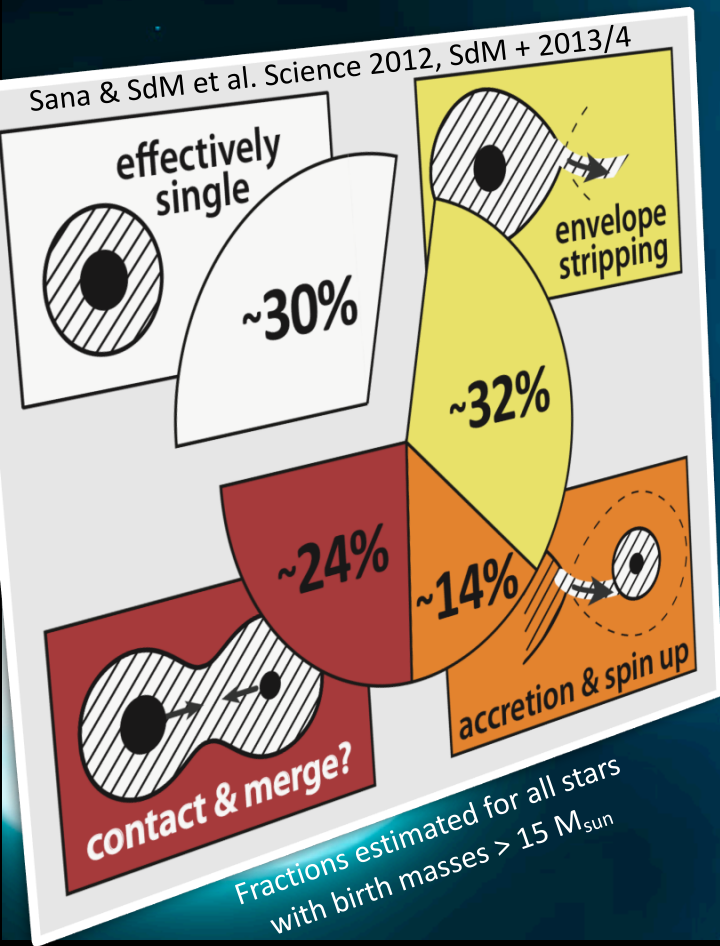




Part I



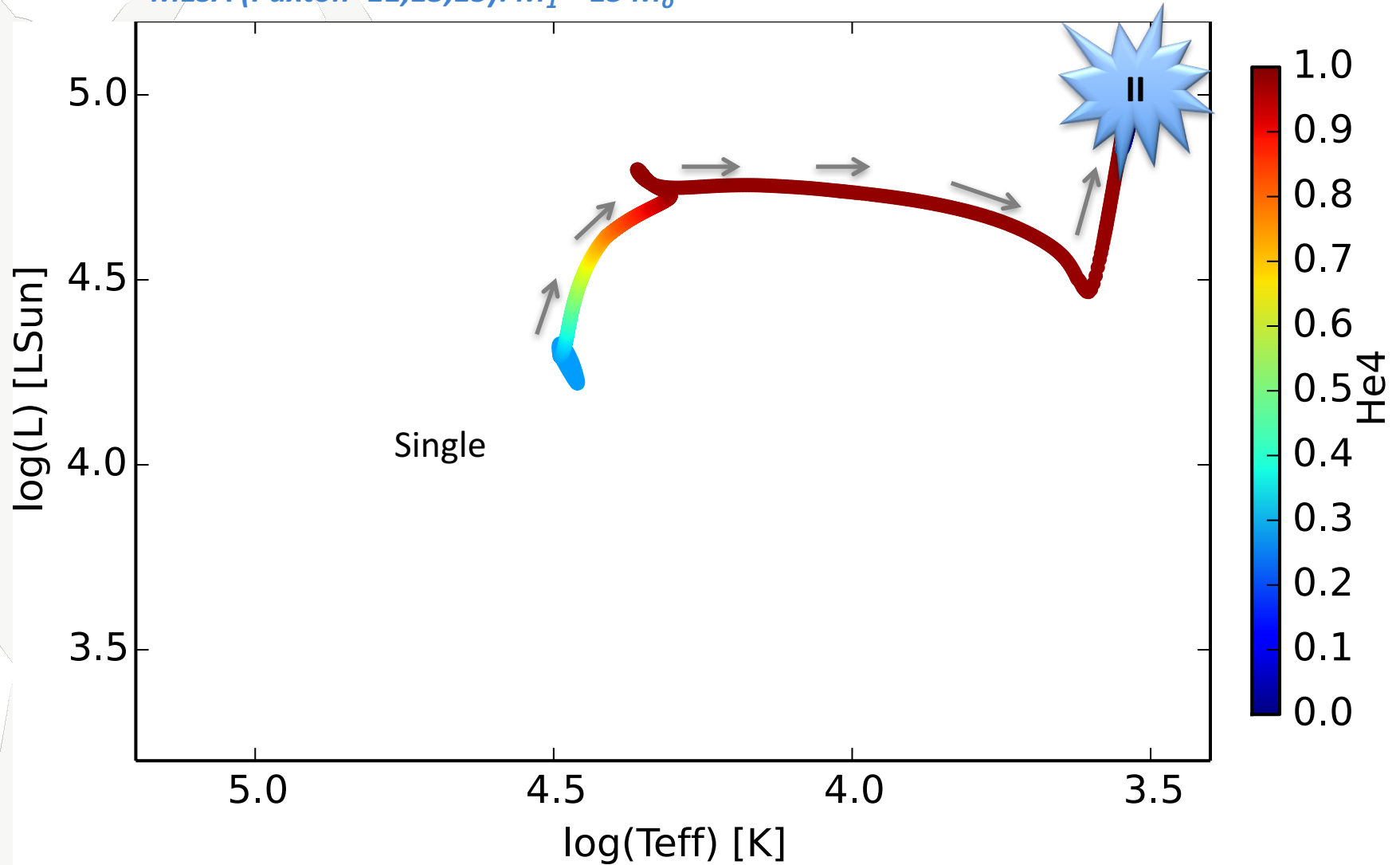
This is not exotic



cf. Abt+78, Kobulnicky+Fryer07, Mason+09,
Chini+12, Sana,SdM+12, Sana,deKoter,SdM+13, Kobulnicky+14,
Dunstall+w/SdM15, Moe+16, Almeida+w/SdM17, ... Moe+18

Evolution of a single star

MESA (Paxton+11,13,15): $M_1 = 15 M_\odot$

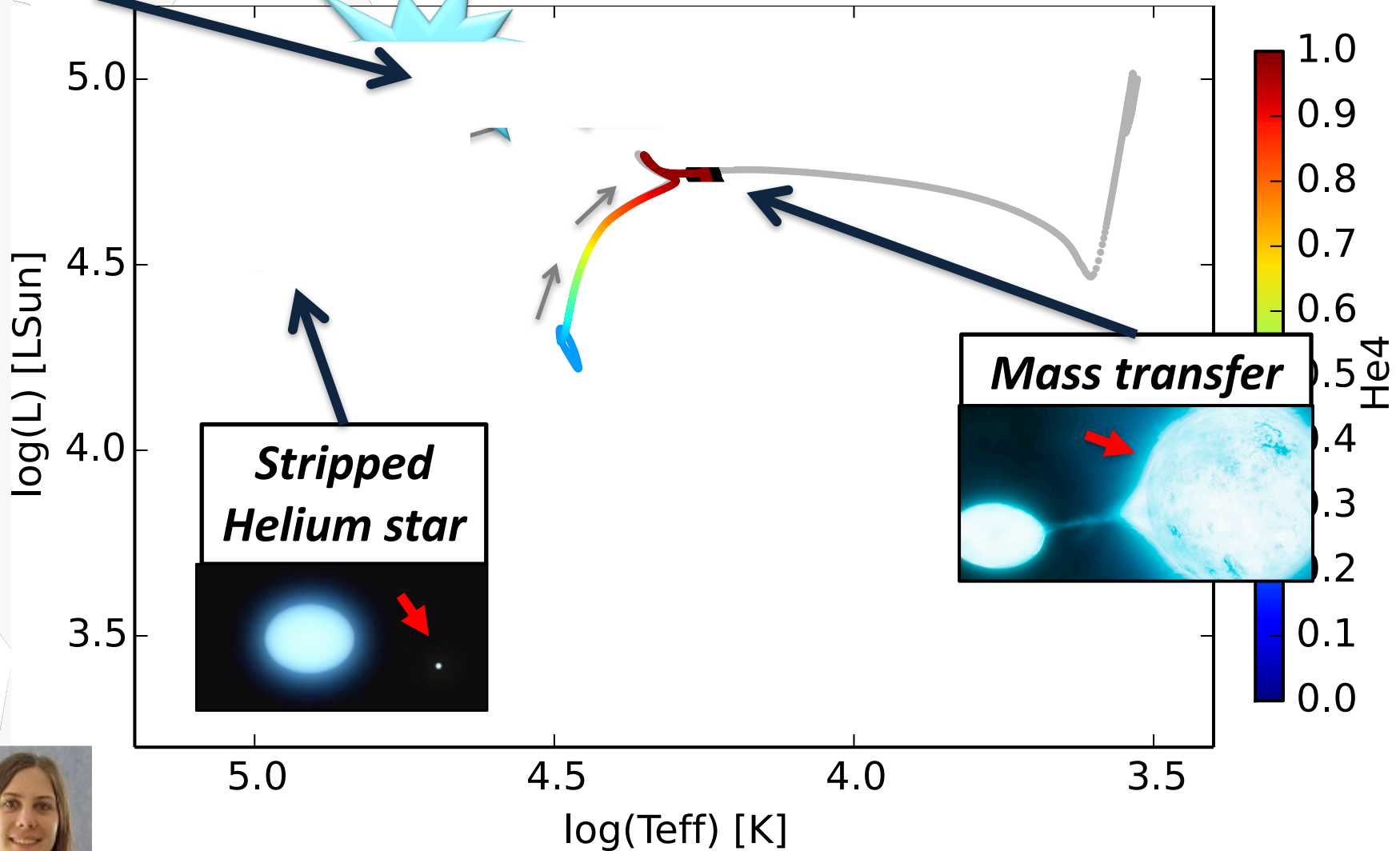


Evolution of a very typical donor star

Götberg et al. (2017),
Laplace et al. (in prep)

He Giant

$M_1 = 15 M_\odot$ $q = 0.8$ *Orbital Period: 10 d*



cf. Kippenhahn, Podsiadlowski, Langer, Wellstein, Pols, Yoon, Claeys, Eldridge, Bersten ...



Eva Laplace

Animation of the life of a typical massive Star



www.eso.org



Part II

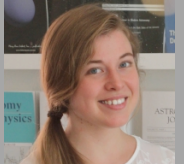
Implications



Applications / Implications

1. Ionizing photons

Götberg et al. 2017, 2018 & to be subm.



**Ylva
Götberg**

2. Late Supernovae

Zapartas et al. 2017, 2018 & to be subm



**Manos
Zapartas**

3. Walkaways & Runaways

Renzo et al. 2018a, b



**Mathieu
Renzo**



New grids of atmosphere models for stripped stars

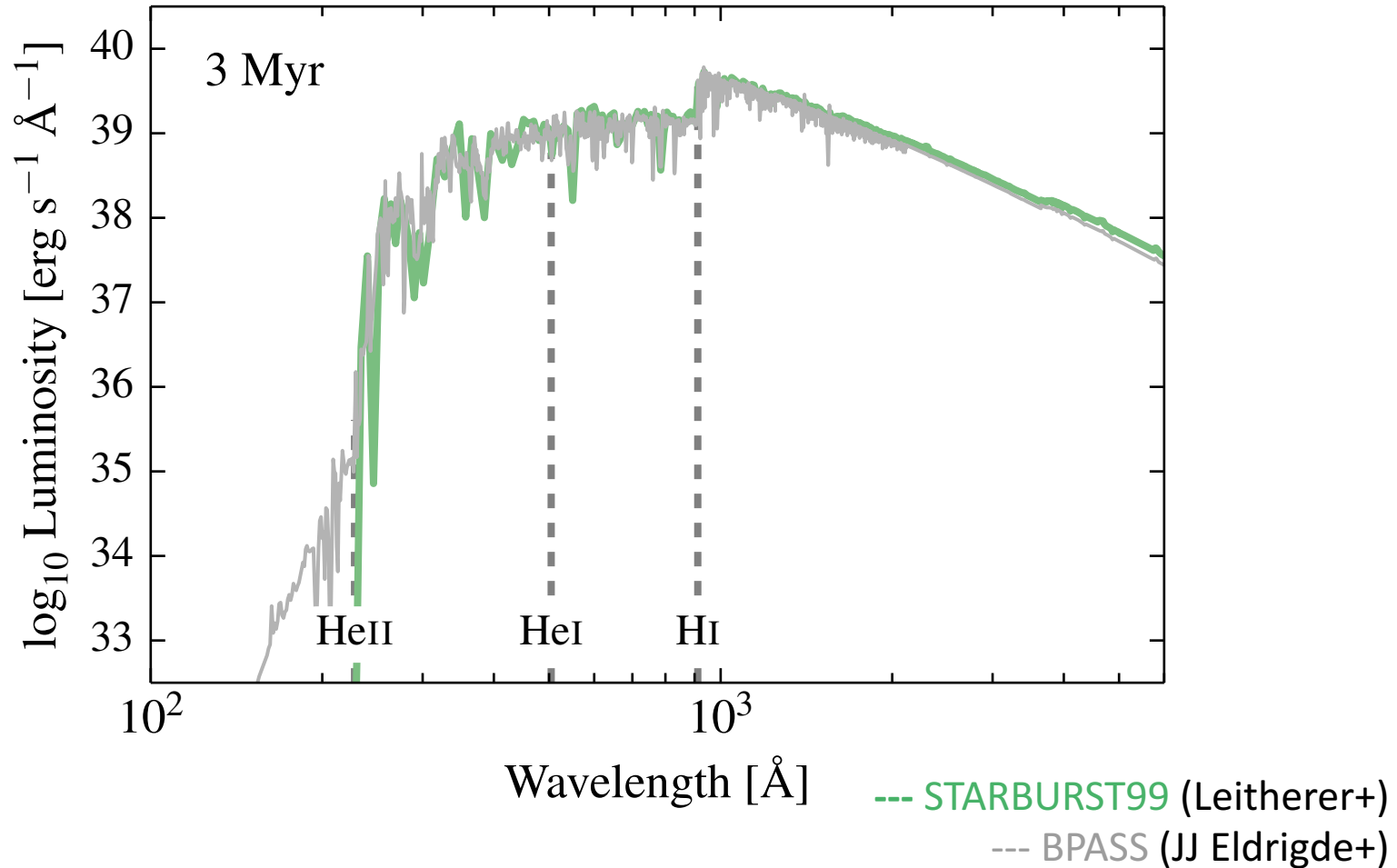
Götberg, de Mink & Groh (2017)
Götberg et al. (2018)



Ylva
Götberg

Implications for integrated spectrum of a Stellar Population

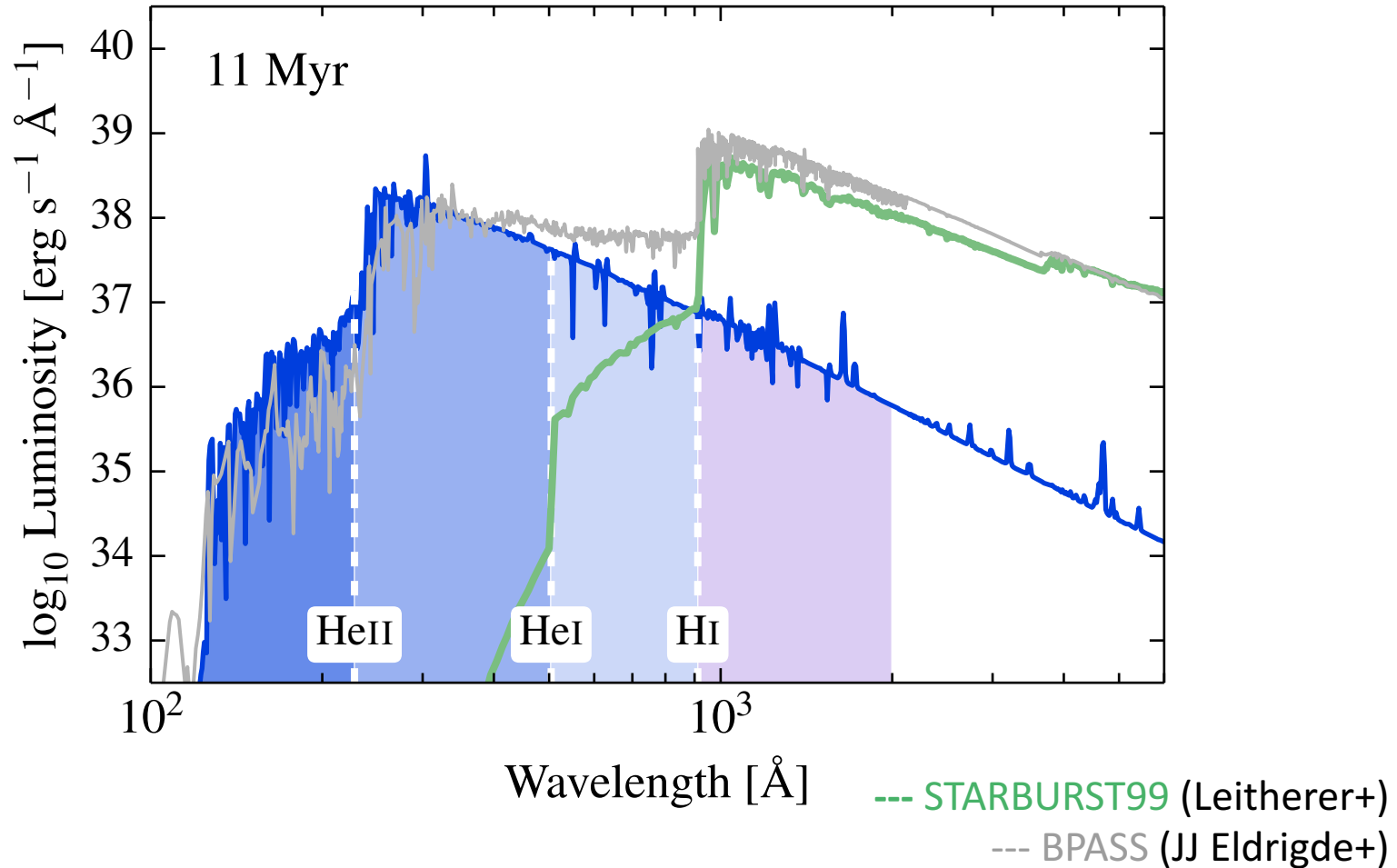
Götberg, dM, Leitherer et al. (to be subm.)



Ylva
Götberg

Implications for integrated spectrum of a Stellar Population

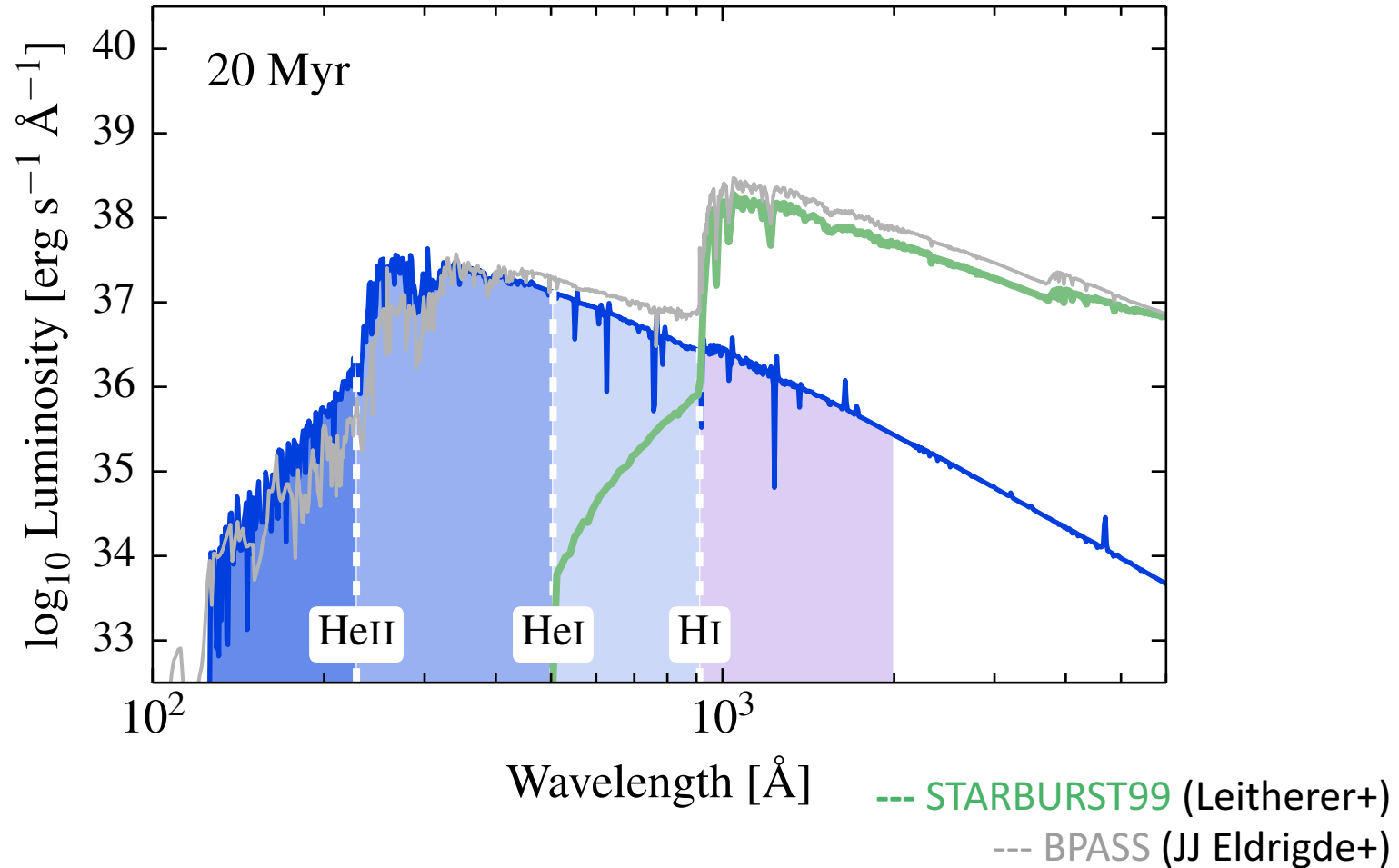
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Ylva
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Implications for integrated spectrum of a Stellar Population

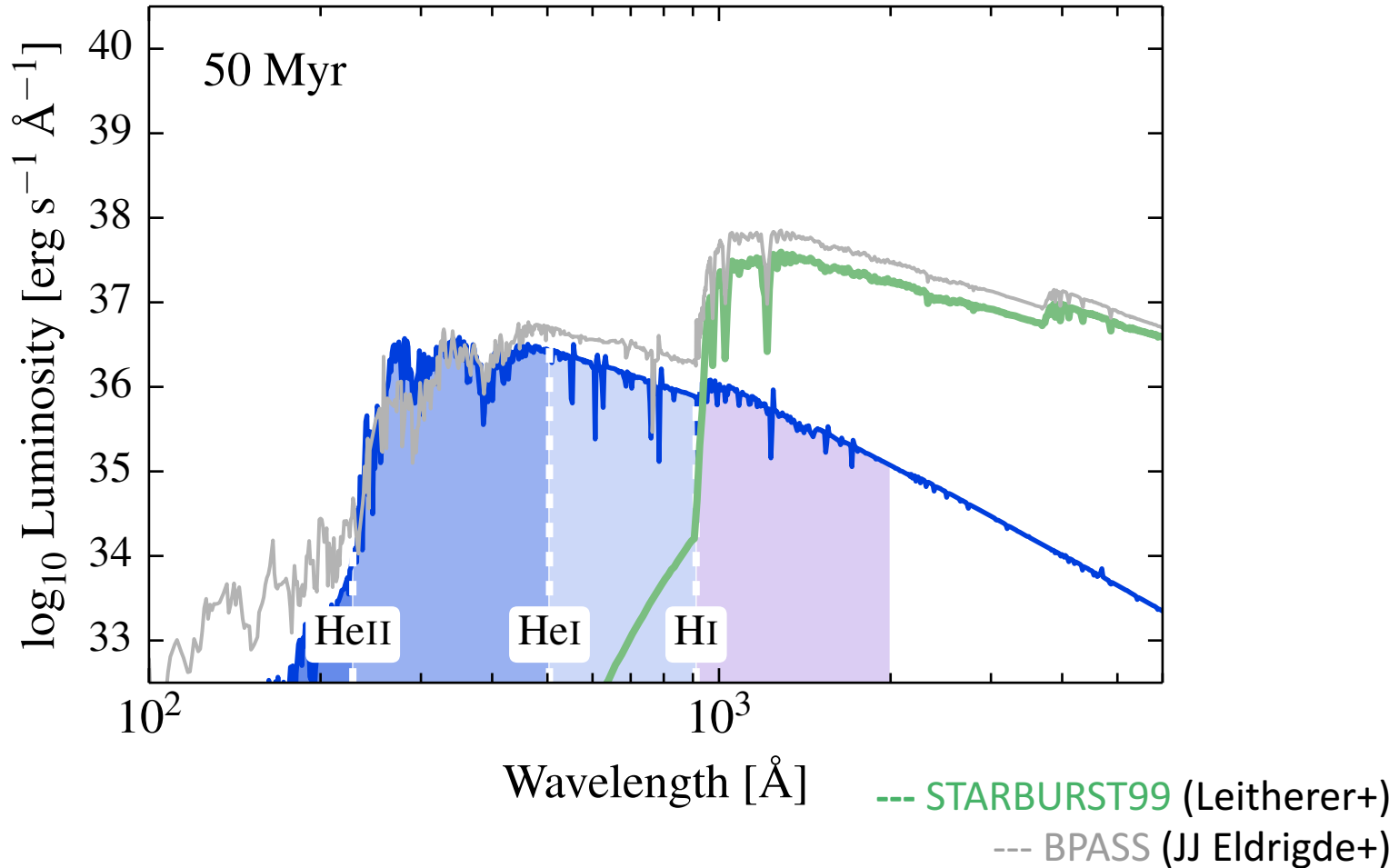
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Ylva
Götberg

Implications for integrated spectrum of a Stellar Population

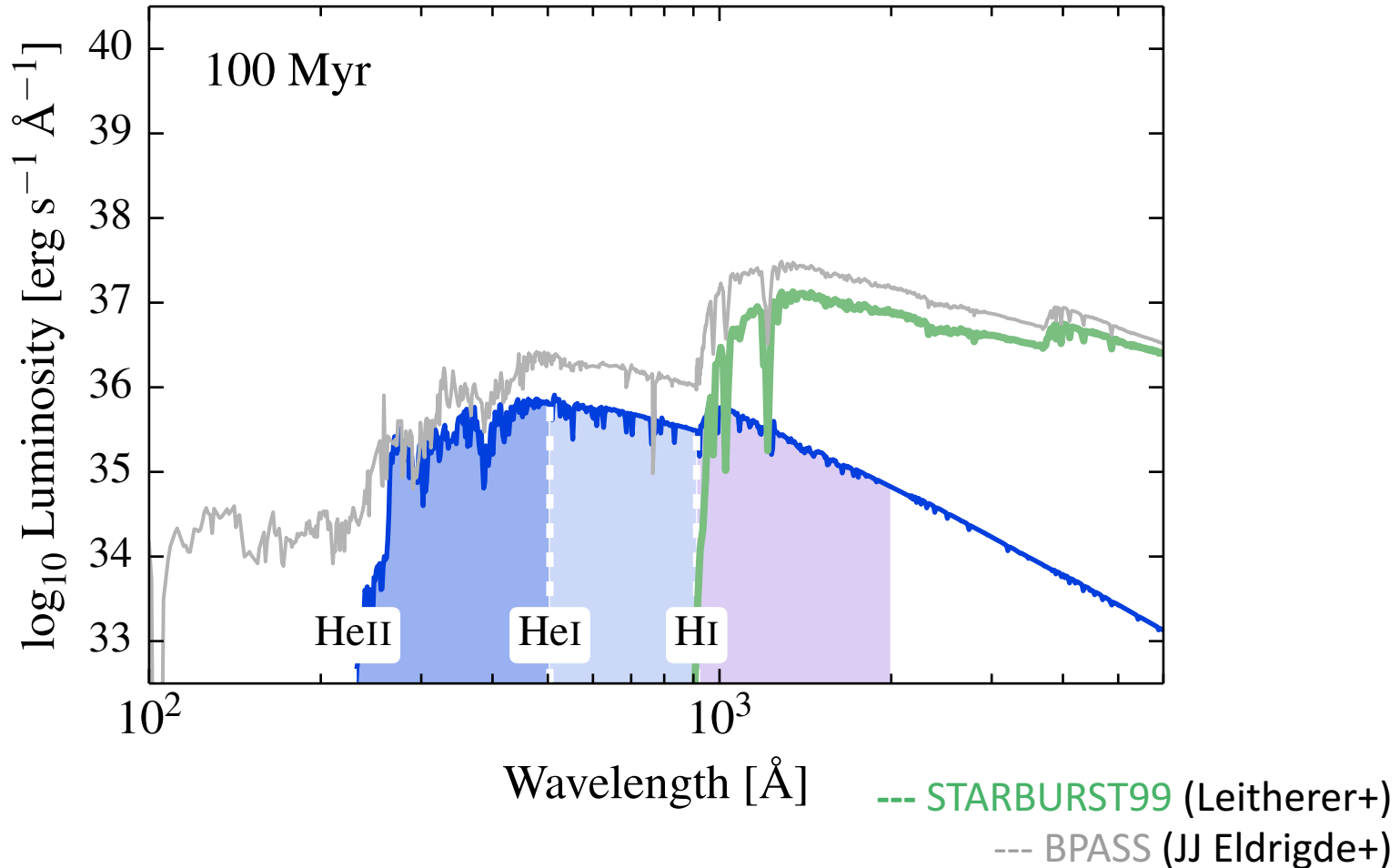
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Ylva
Götberg

Implications for integrated spectrum of a Stellar Population

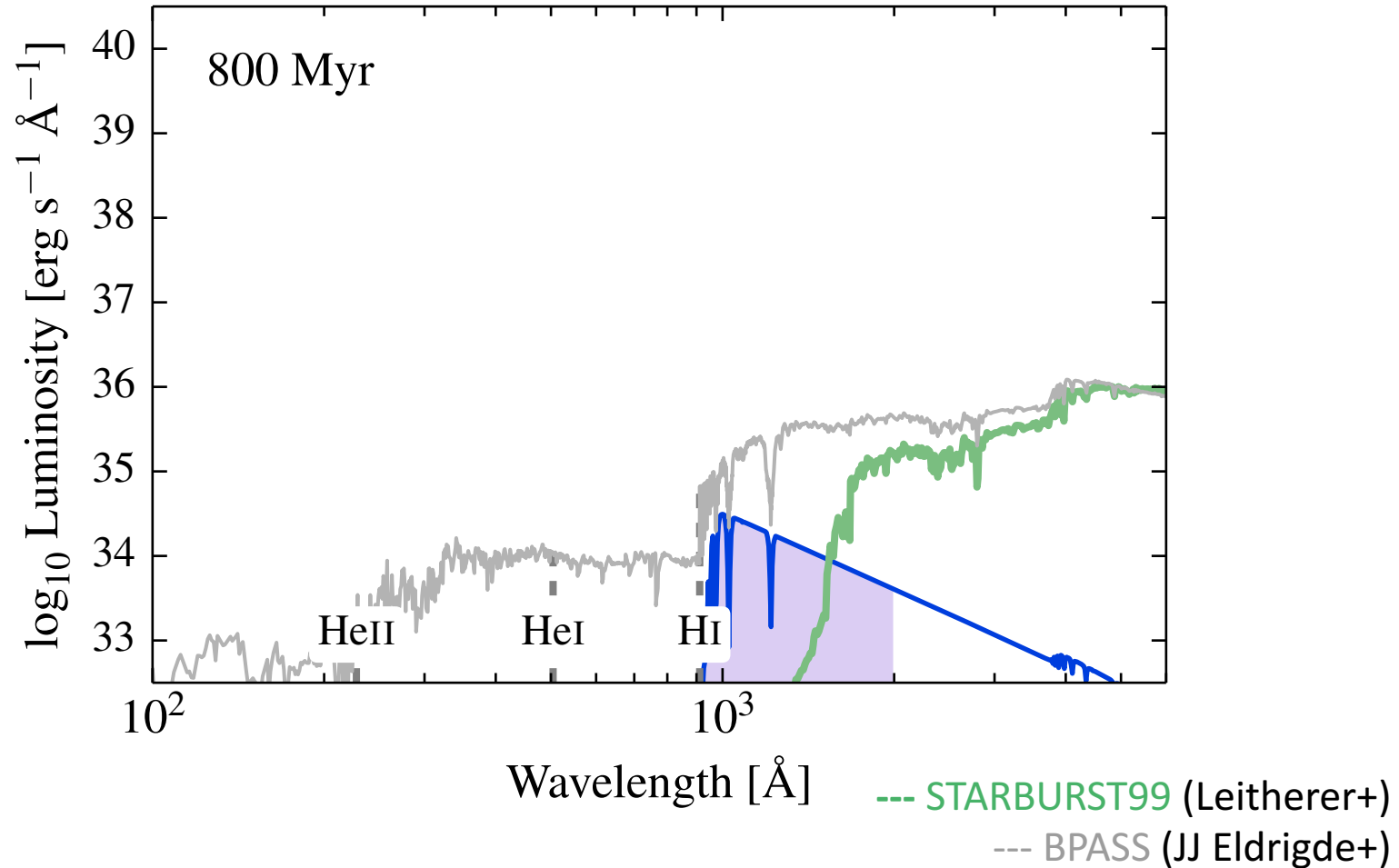
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Ylva
Götberg

Implications for integrated spectrum of a Stellar Population

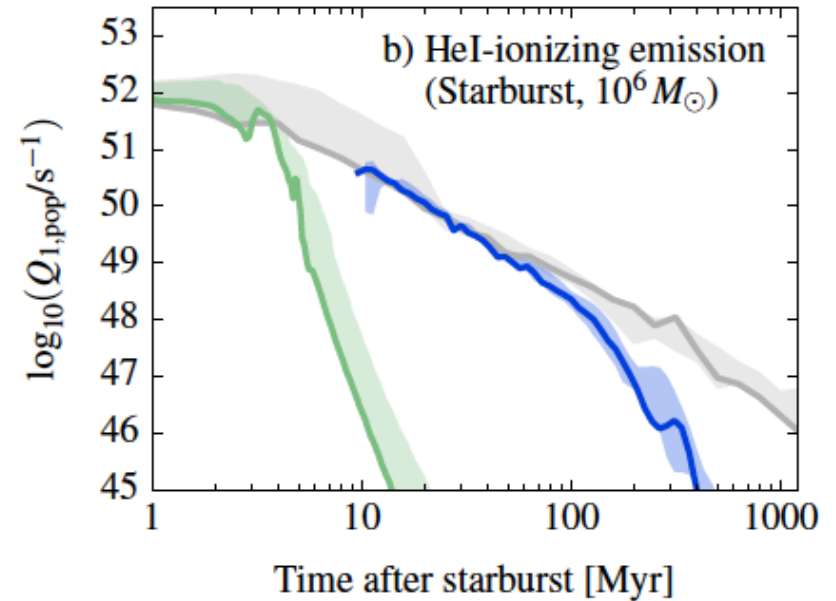
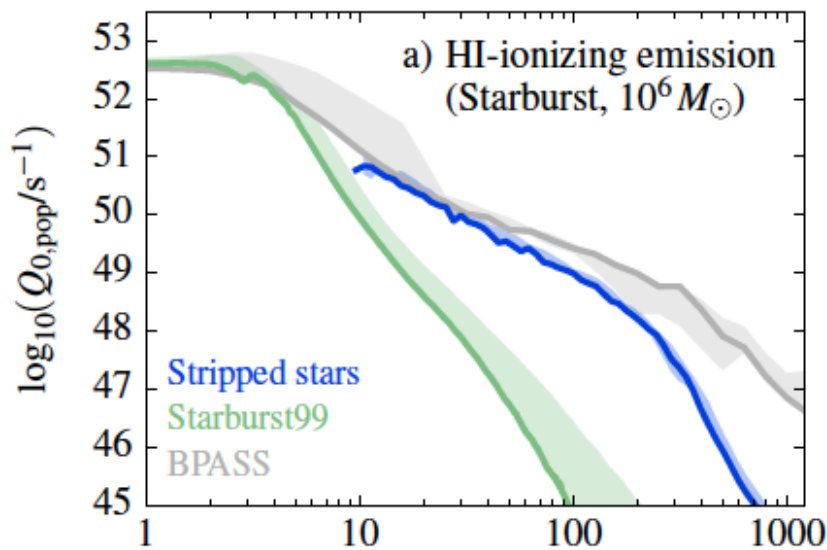
Götberg, dM, Leitherer et al. (to be subm.)



Ylva
Götberg

Ionizing photons

Götberg, dM, Leitherer et al. (to be subm.)



Ylva
Götberg

Possible contribution to HI - Reionization

Götberg, dM, et al. (in prep.)

**Warning
preliminary**



Ylva
Götberg



Applications / Implications

1. Ionizing photons

Götberg et al. 2017, 2018 & to be subm.

2. Late Supernovae

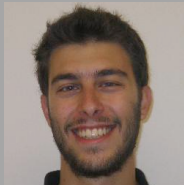
Zapartas et al. 2017, 2018 & to be subm

3. Walkaways & Runaways

Renzo et al. 2018a, b



**Ylva
Götberg**



**Manos
Zapartas**



**Mathieu
Renzo**



I. When?

Delay time distribution:

“Supernova rate”

First massive star dies
~3 Myr

Last massive star dies:
~ 48 Myr



Manos Zapartas

birth

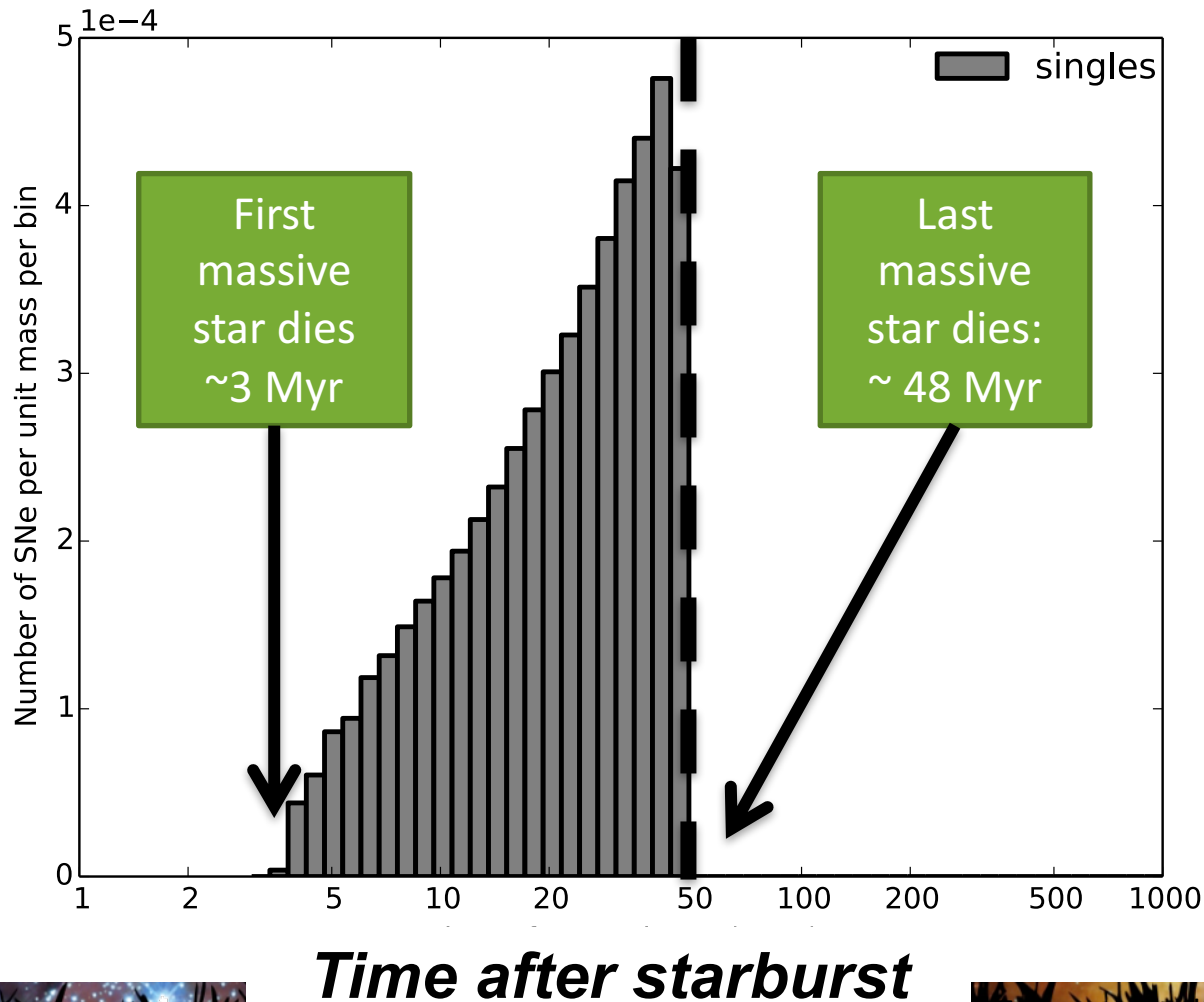
Time after starburst



Result for single stars

Zapartas & de Mink et al. (2017,8)

“Supernova rate”



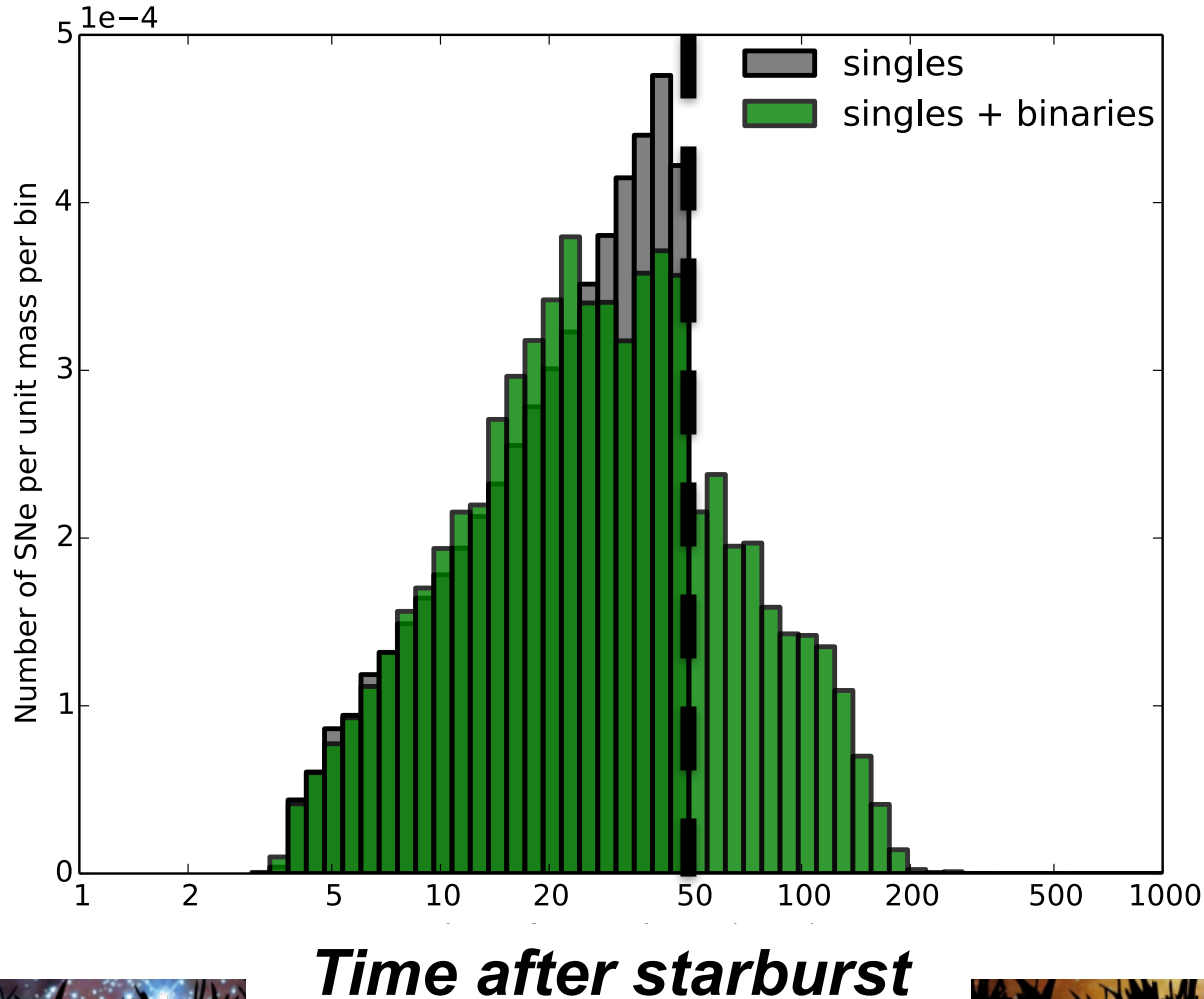
Manos Zapartas

Including binaries

Compare with work by Vanbeveren+ and Eldridge+
Simulations with Rob Izzard's Binary_c code

Zapartas & de Mink et al. (2017,8)

“Supernova rate”



Manos Zapartas



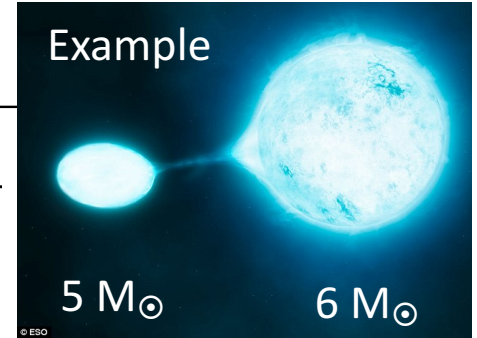
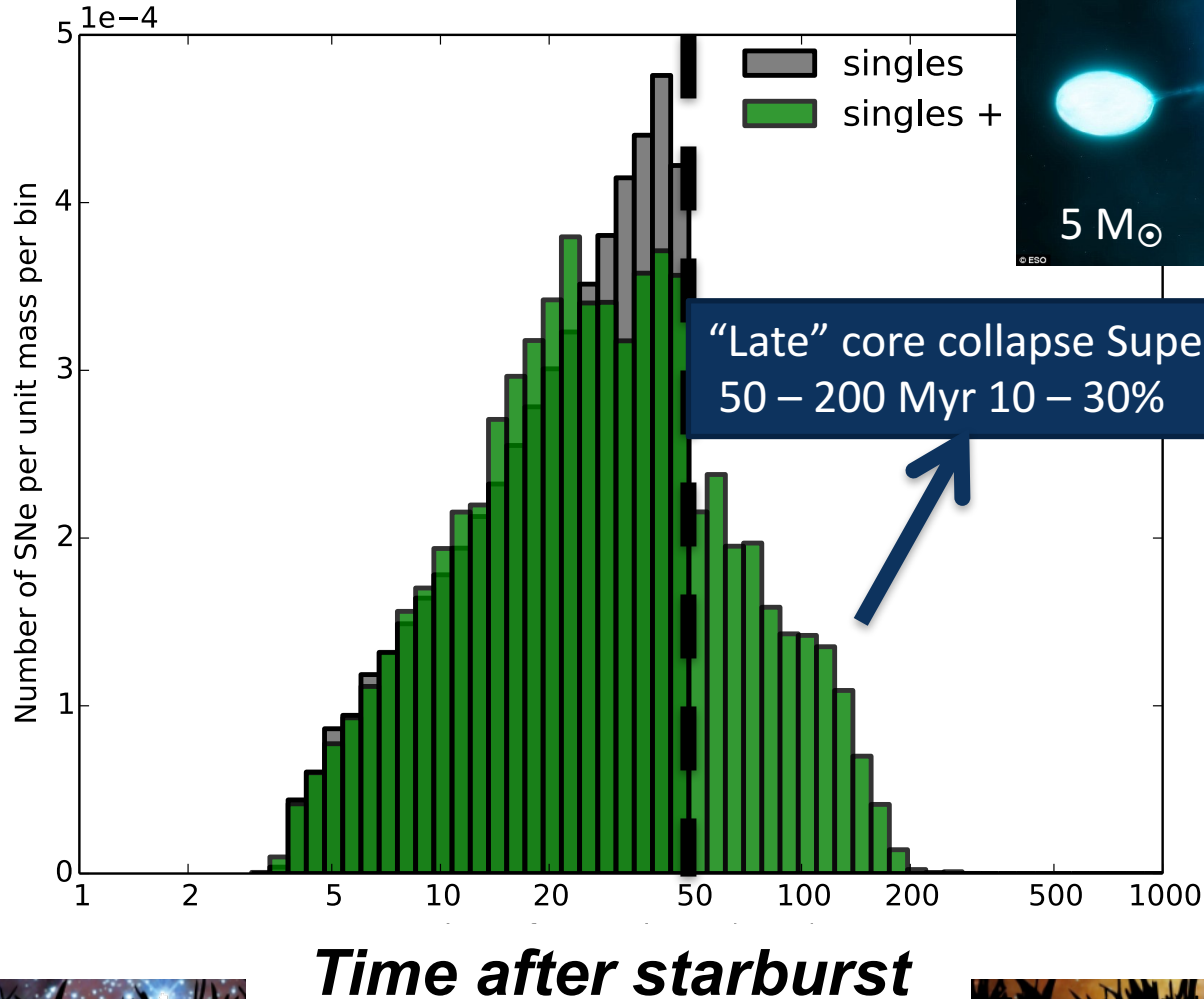
Including binaries

cf. Vanbeveren+.and Eldridge+
Simulations with Rob Izzard's Binary_c code

Zapartas & de Mink et al. (2017, 8)

1. More supernovae: up to ~30%

“Supernova rate”



“Late” core collapse Supernovae:
50 – 200 Myr 10 – 30%

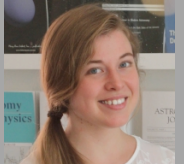


Manos
Zapartas

Applications / Implications

1. Ionizing photons

Götberg et al. 2017, 2018 & to be subm.



**Ylva
Götberg**

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**Manos
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Renzo et al. 2018a, b



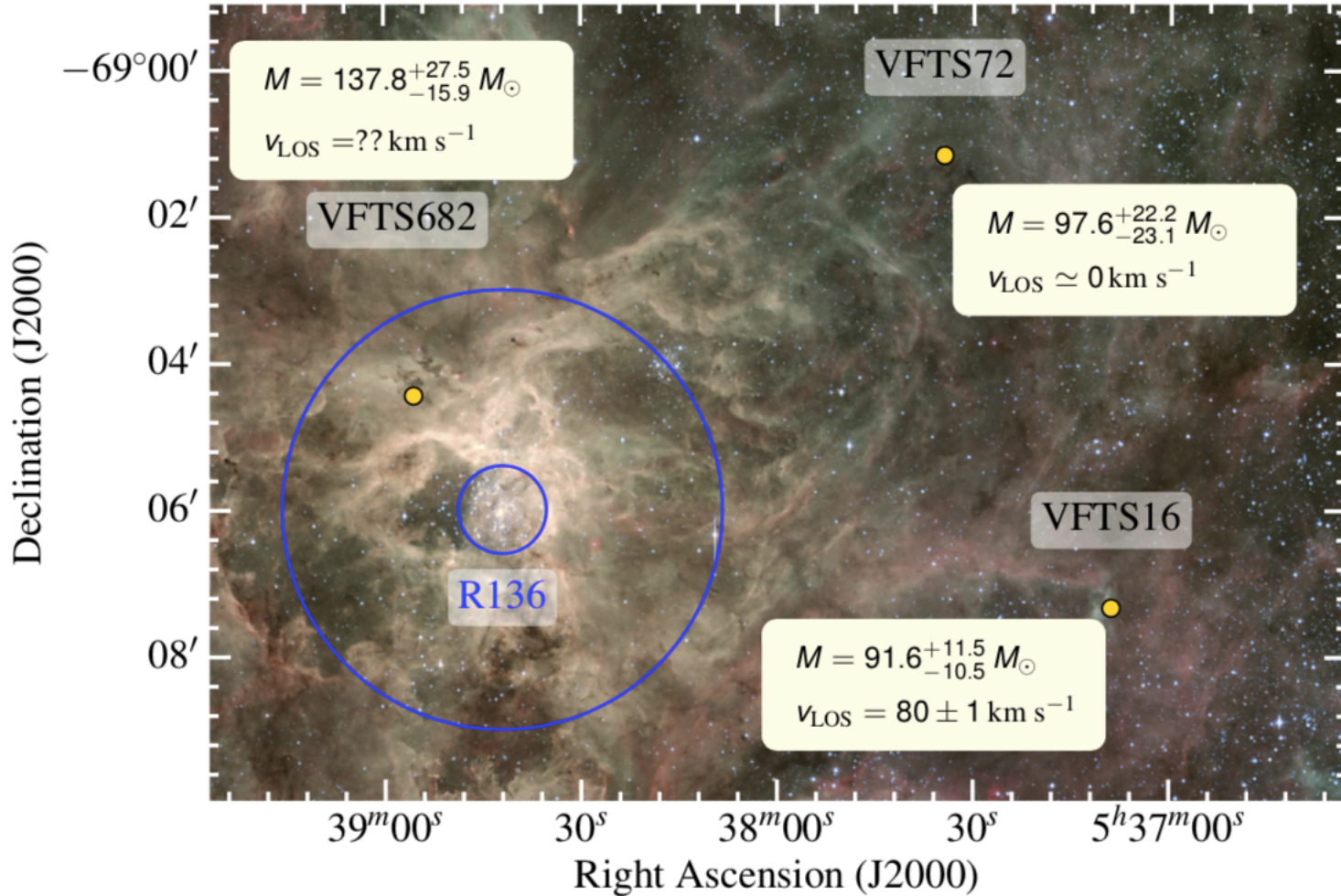
**Mathieu
Renzo**





Mathieu Renzo

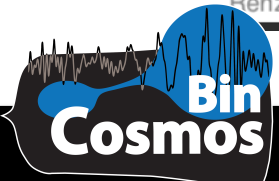
The most massive runaways known



Renzo *et al.*, submitted

Lennon *et al.*, accepted, arXiv:1805.08277 5/17

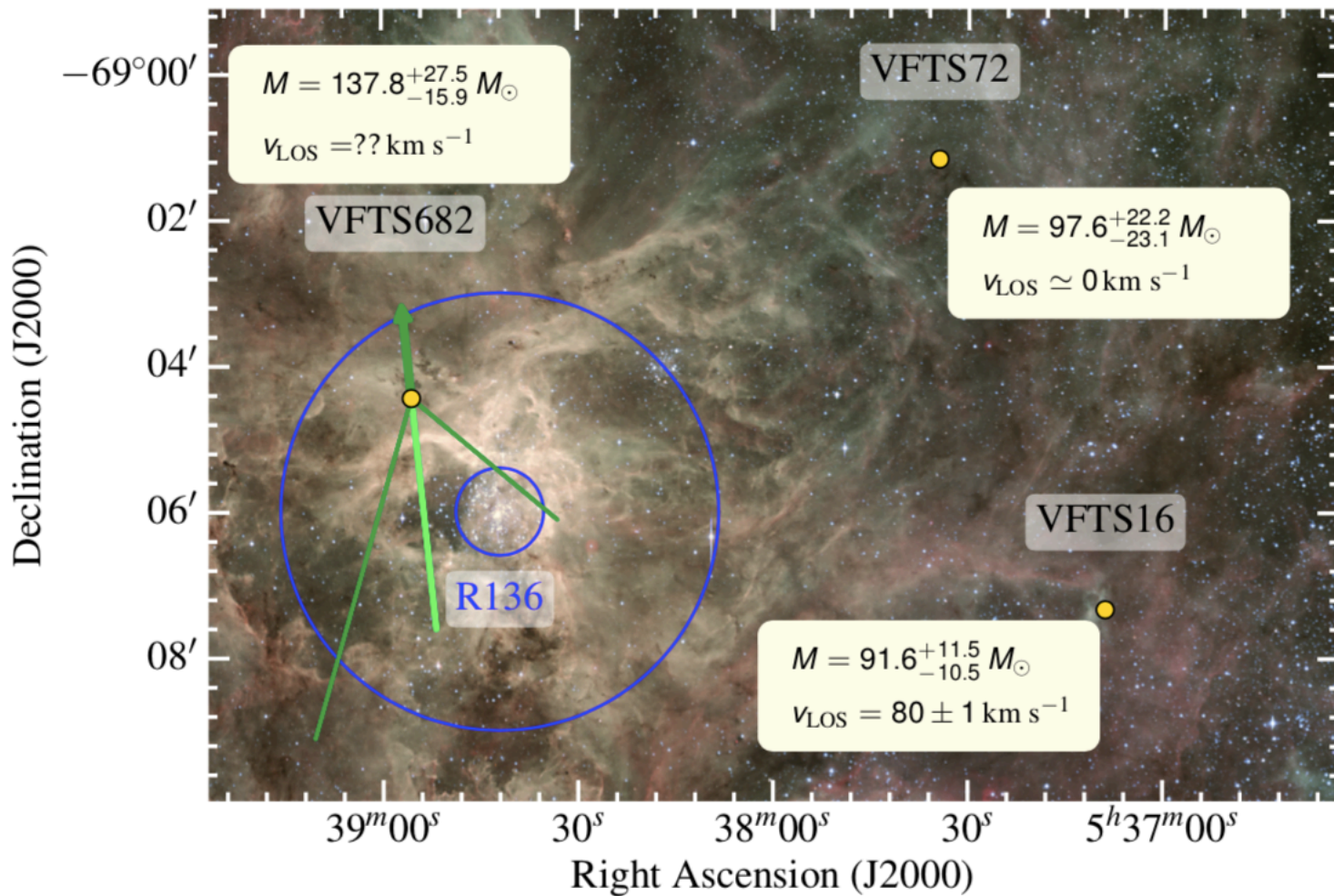
Renzo *et al.* 2018a, b





Mathieu Renzo

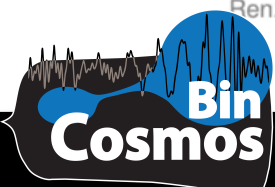
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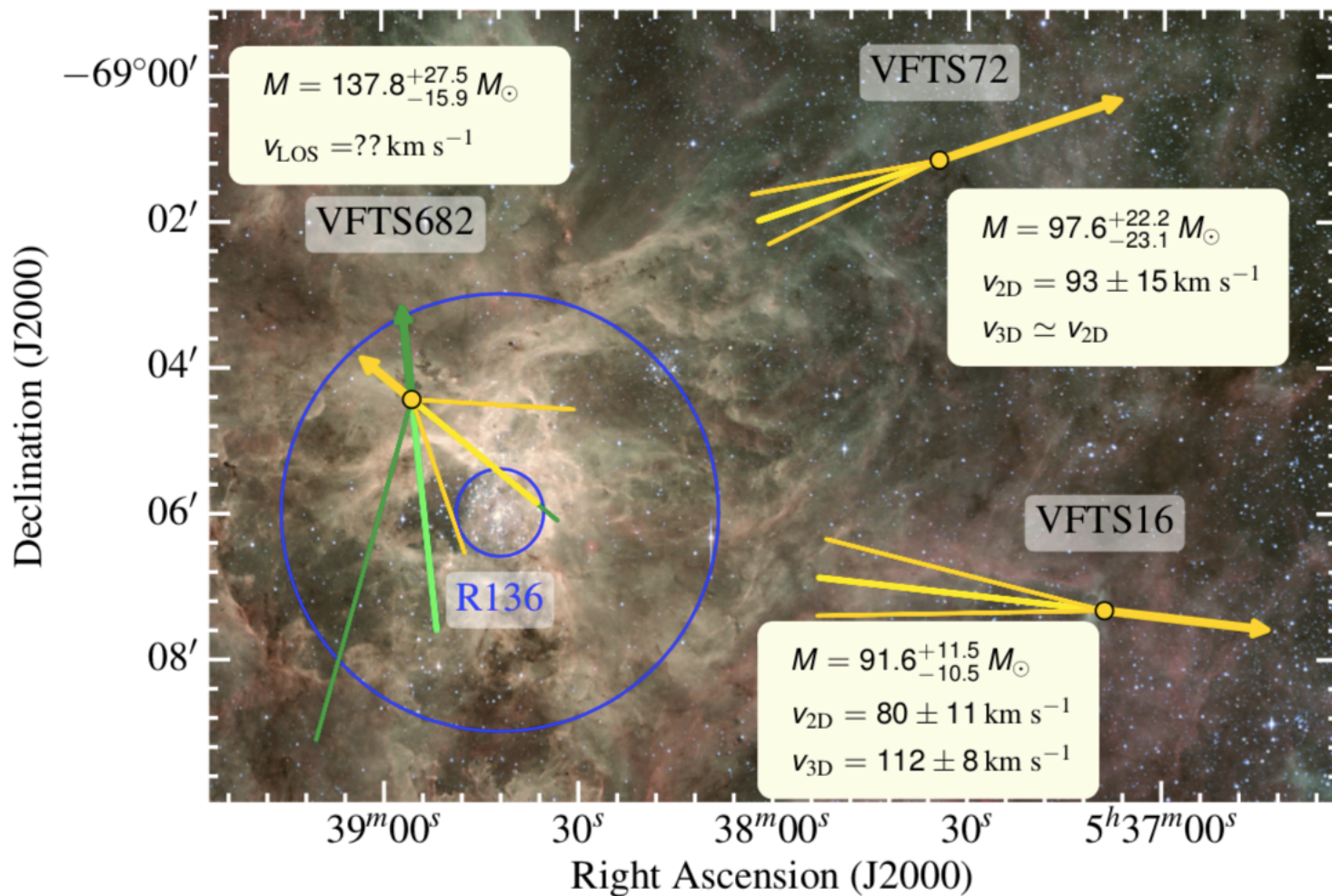
Renzo *et al.* 2018a, b





Mathieu Renzo

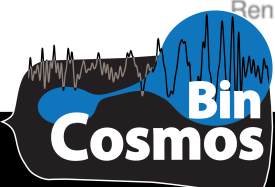
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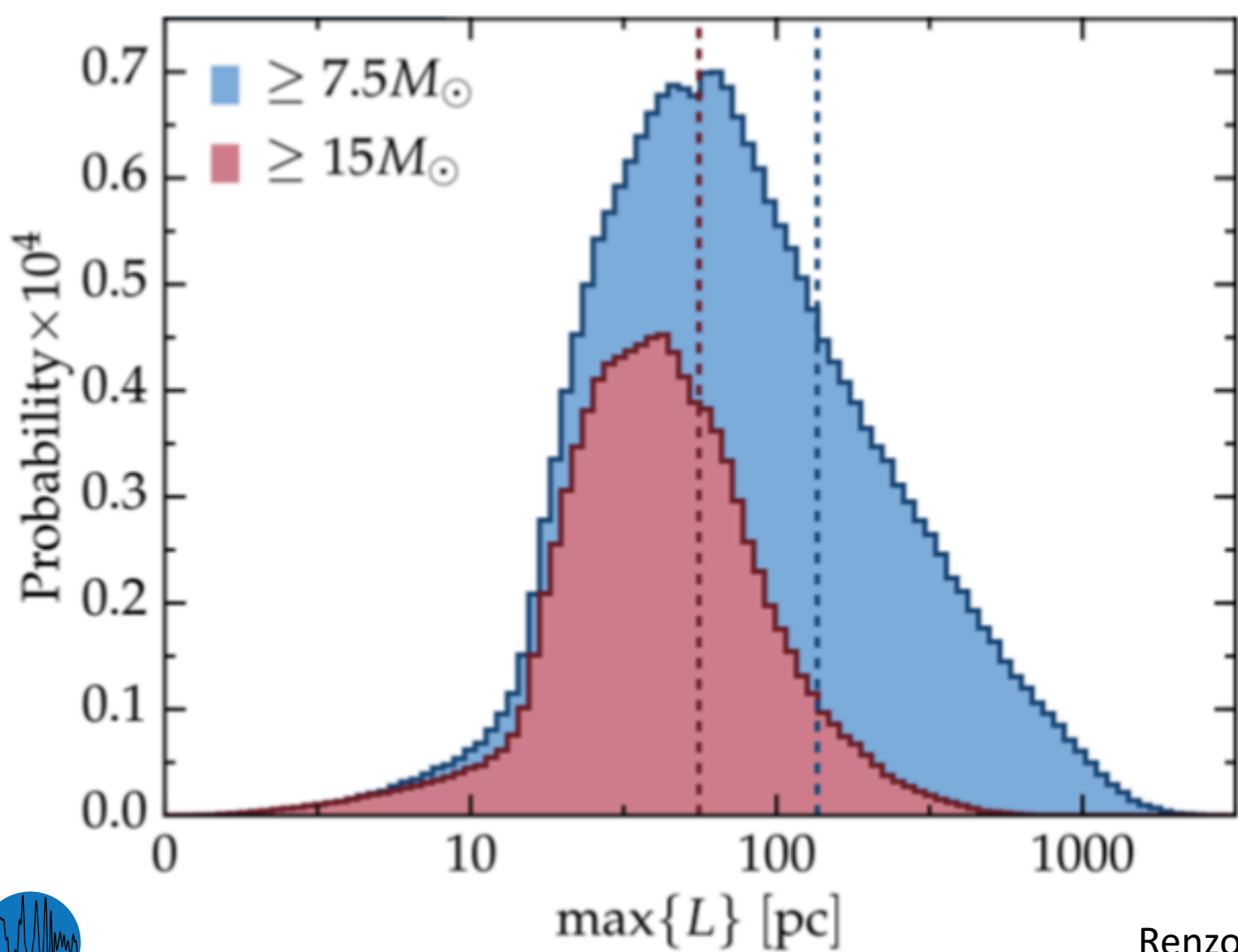
Renzo et al. 2018a, b



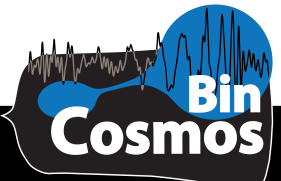


Mathieu Renzo

How far do they get?



Renzo et al. 2018a, b





Conclusions



How complete is this picture?

Stars > 100 Msun?

X-ray binaries

Stripped stars in binaries?

Trends with metallicity?

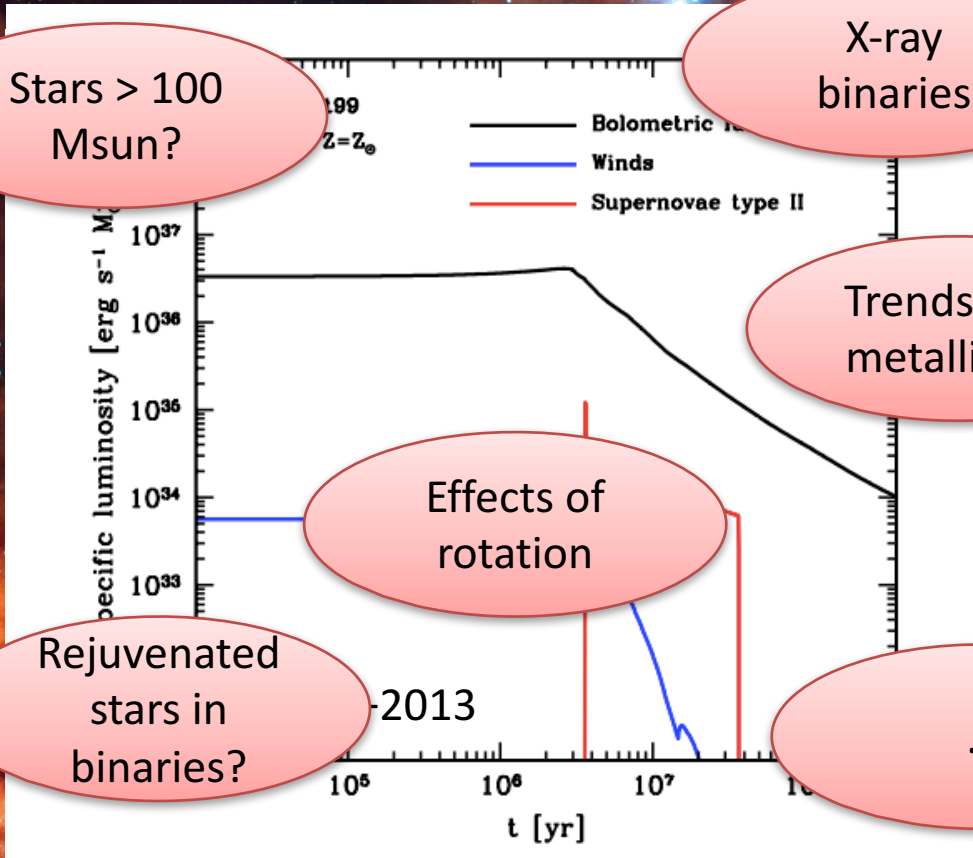
Walkaway / runaways

Effects of rotation

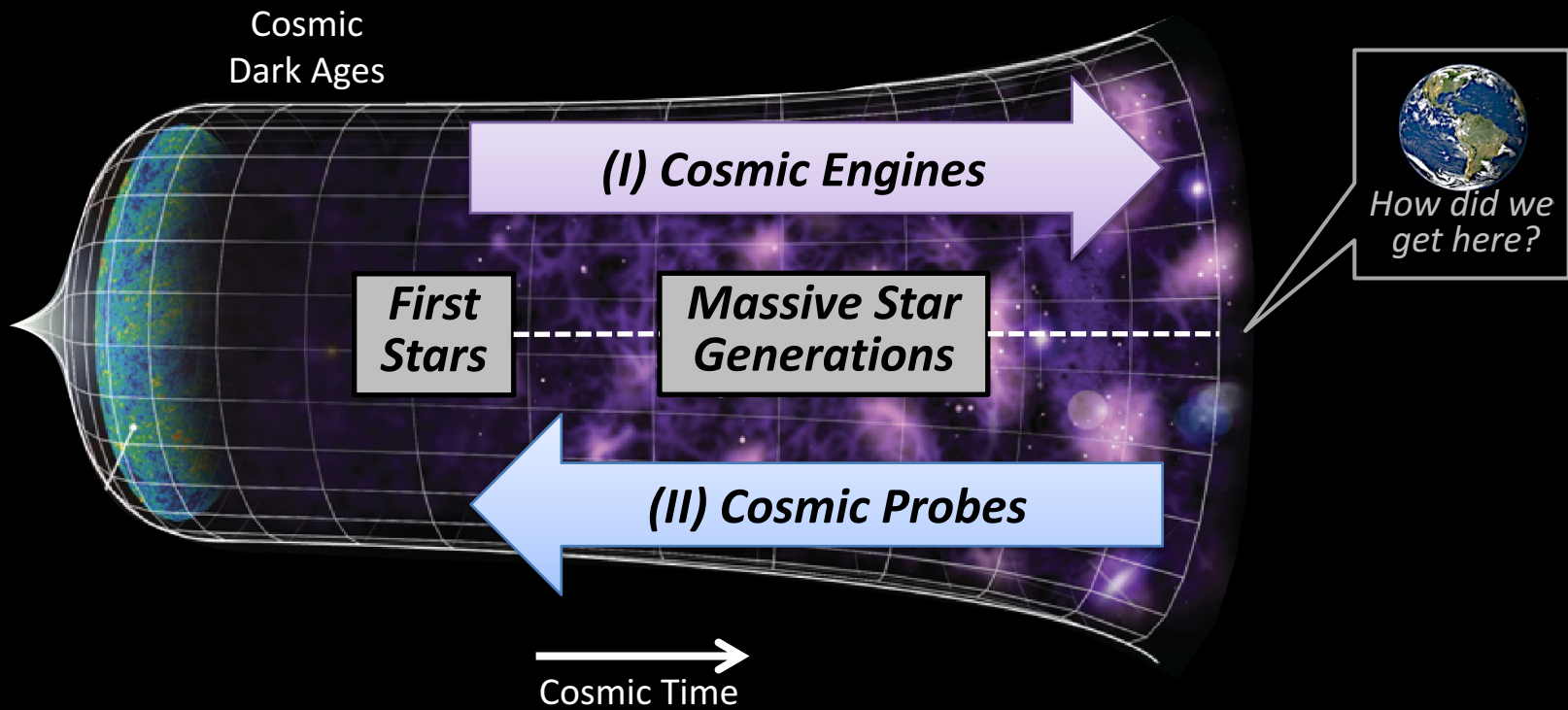
Rejuvenated stars in binaries?

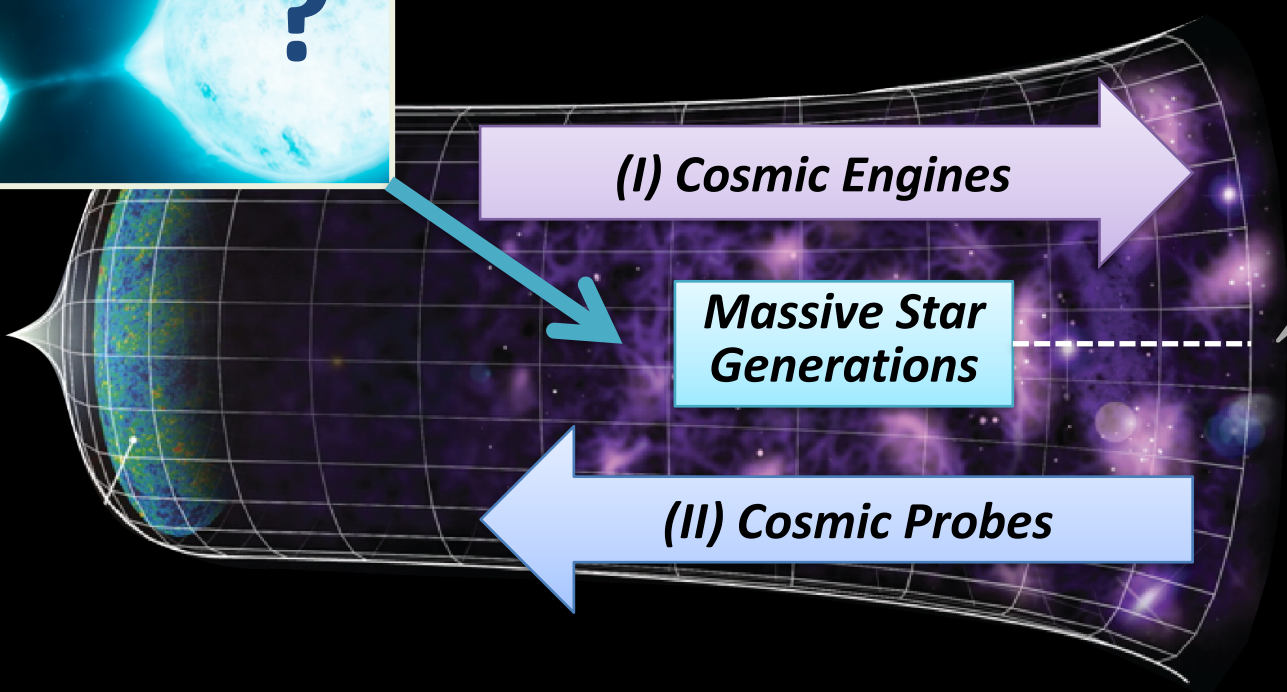
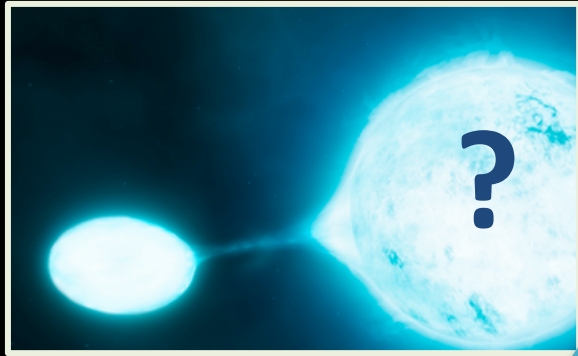
Triples

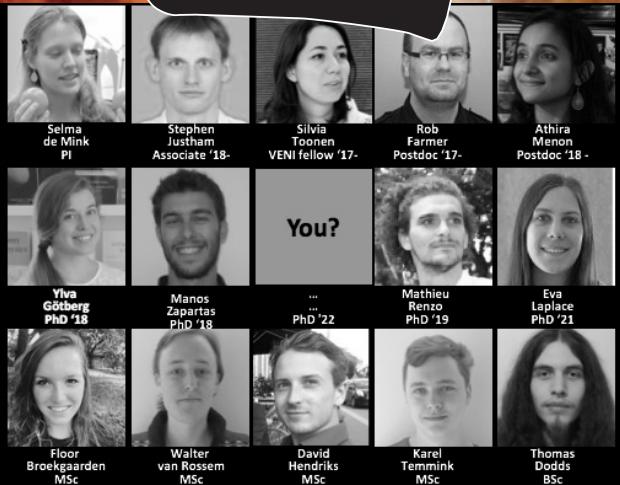
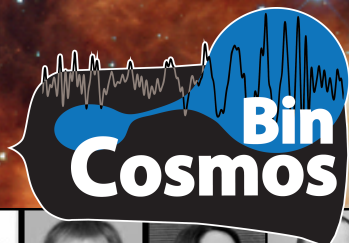
...



Cosmic
Dark Ages



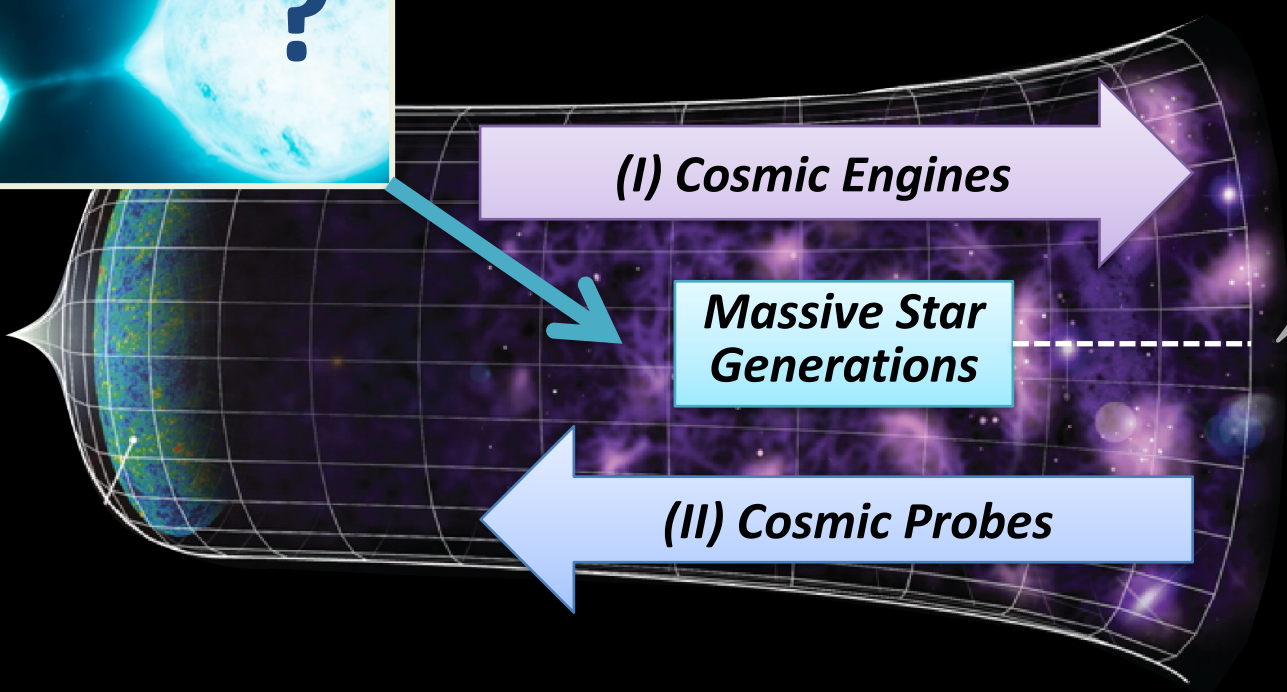
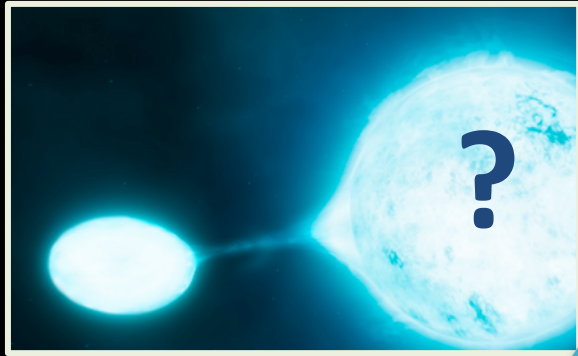




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VLT-FLAMES Massive Star Consortium,



(I) Cosmic Engines

*Massive Star
Generations*

(II) Cosmic Probes



How did we
get here?